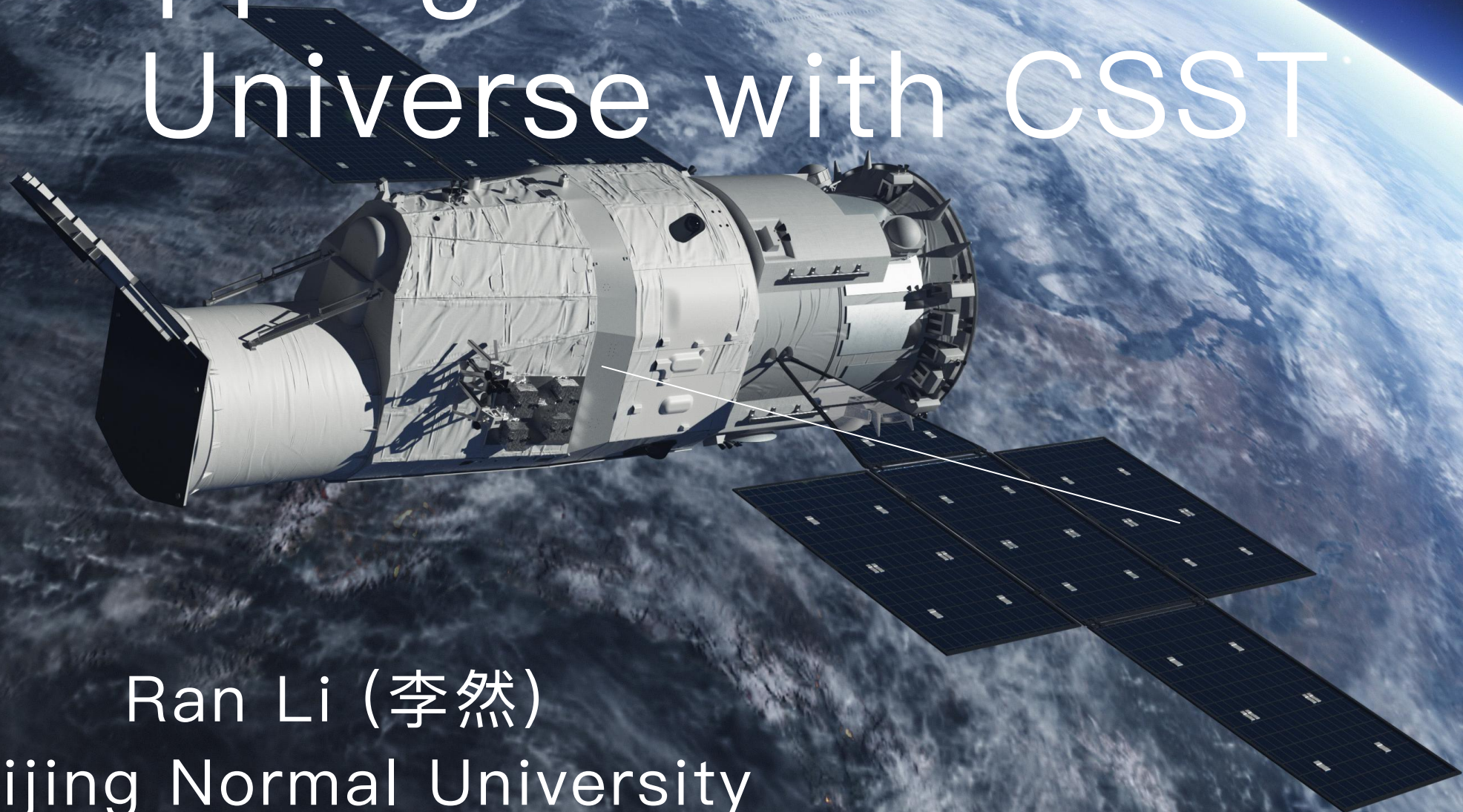


Mapping the Dark Matter Universe with CSST



Ran Li (李然)
Beijing Normal University

Mapping the Dark Matter Universe with CSST

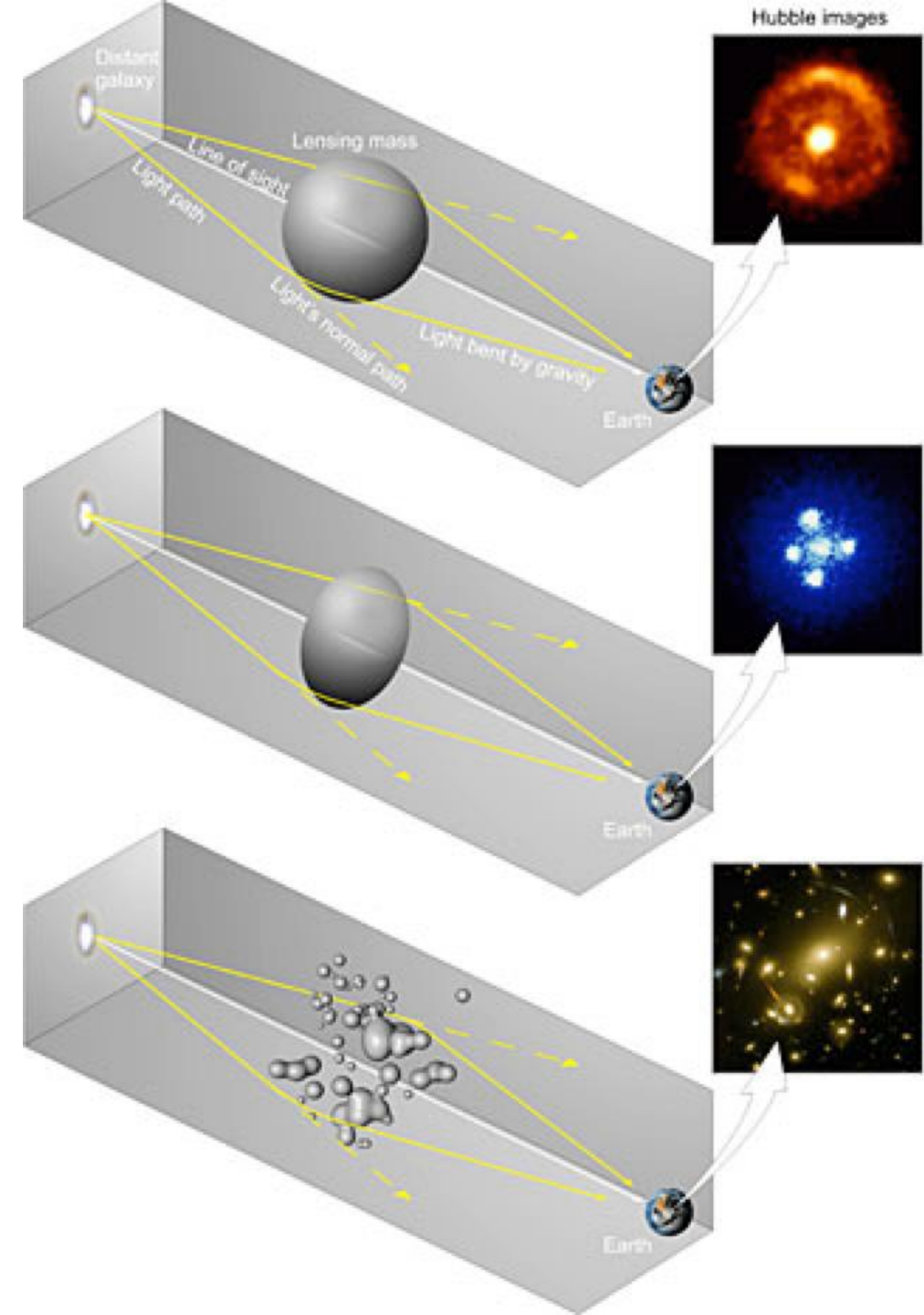
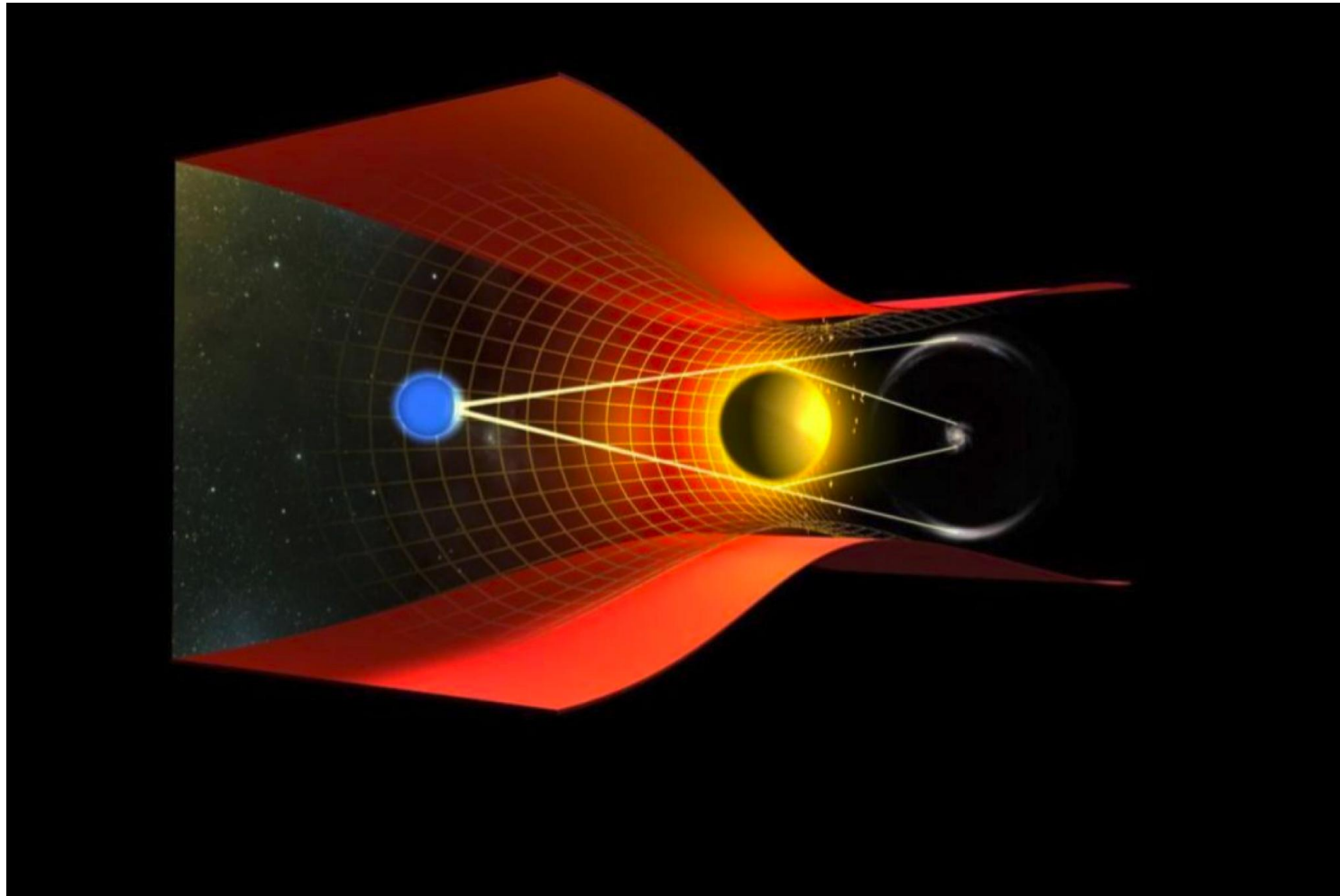
In the Age of AI

Ran Li (李然)
Beijing Normal University

A visualization of the cosmic web, showing a complex network of dark matter filaments and clusters. The background is dark, with numerous bright yellow and orange points representing galaxies and galaxy clusters. The filaments are thin, dark lines that connect these clusters, forming a web-like structure. The overall effect is a sense of vastness and complexity in the universe's structure.

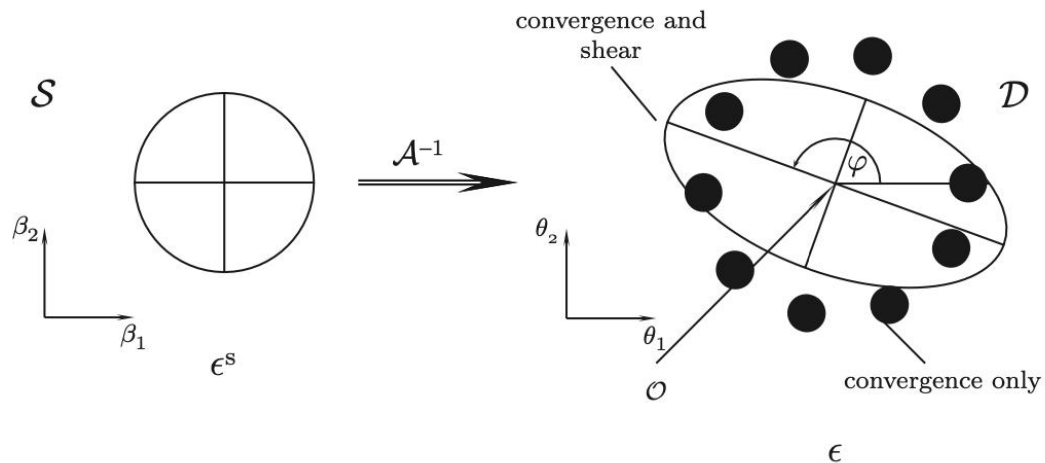
Dark Matter Dominates the Structure Formation and Evolution of Our Universe

Mapping the Dark Matter with Gravitational Lensing

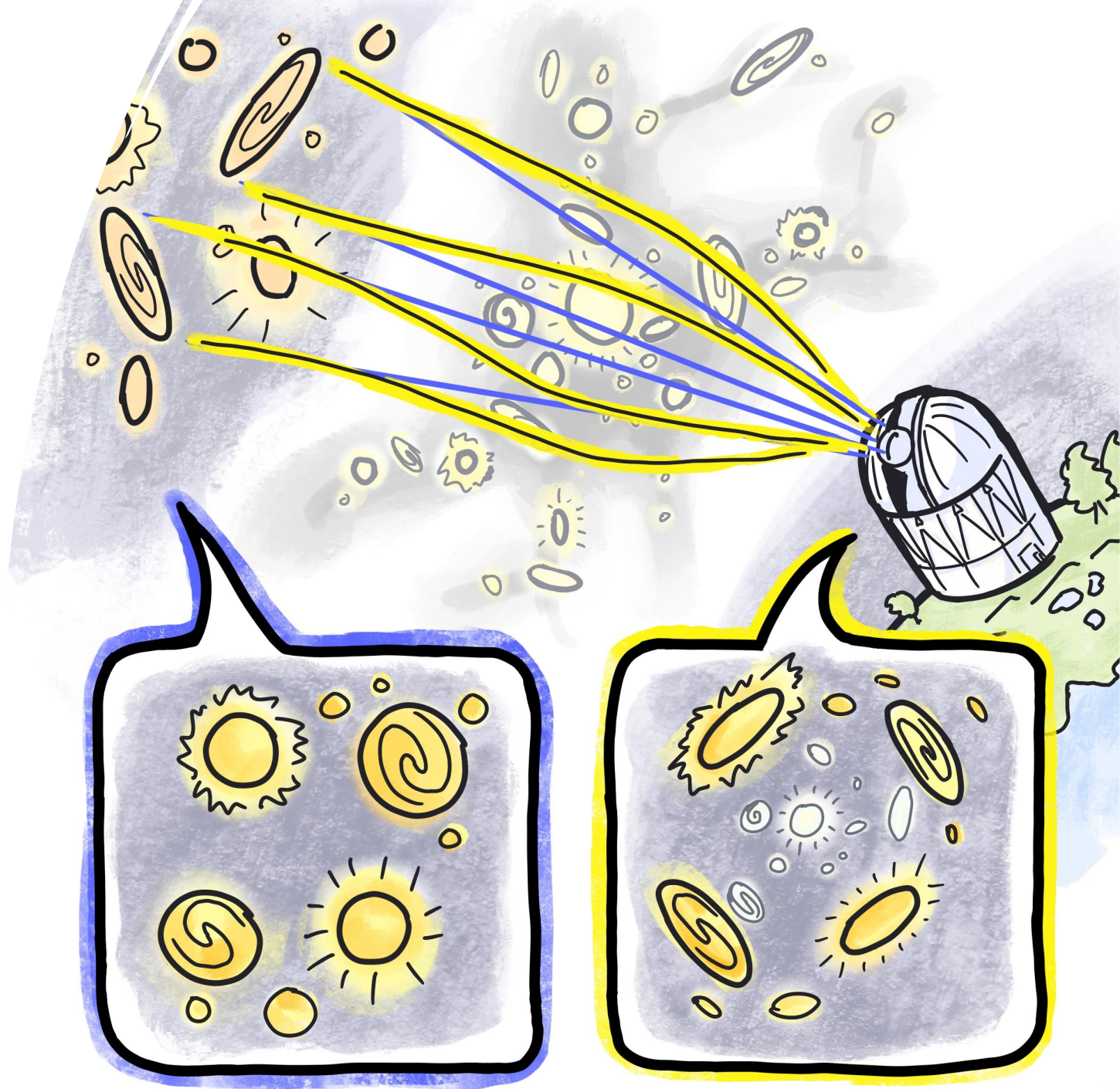


Weak Lensing Shear (γ)

$$\mathcal{A}(\boldsymbol{\theta}) = (1 - \kappa) \begin{pmatrix} 1 - g_1 & -g_2 \\ -g_2 & 1 + g_1 \end{pmatrix}, \text{ where } g(\boldsymbol{\theta}) = \frac{\gamma(\boldsymbol{\theta})}{[1 - \kappa(\boldsymbol{\theta})]}$$



M. Bradač

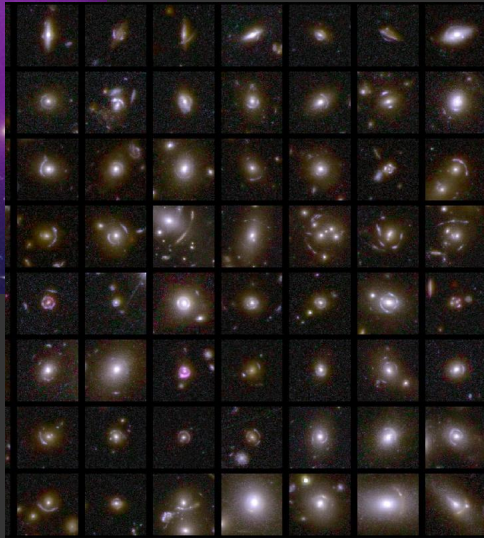


Credit: Jessie Muir 20

– Ongoing and Upcoming Space Optical Surveys



Euclid Telescope
and Q1 Lenses

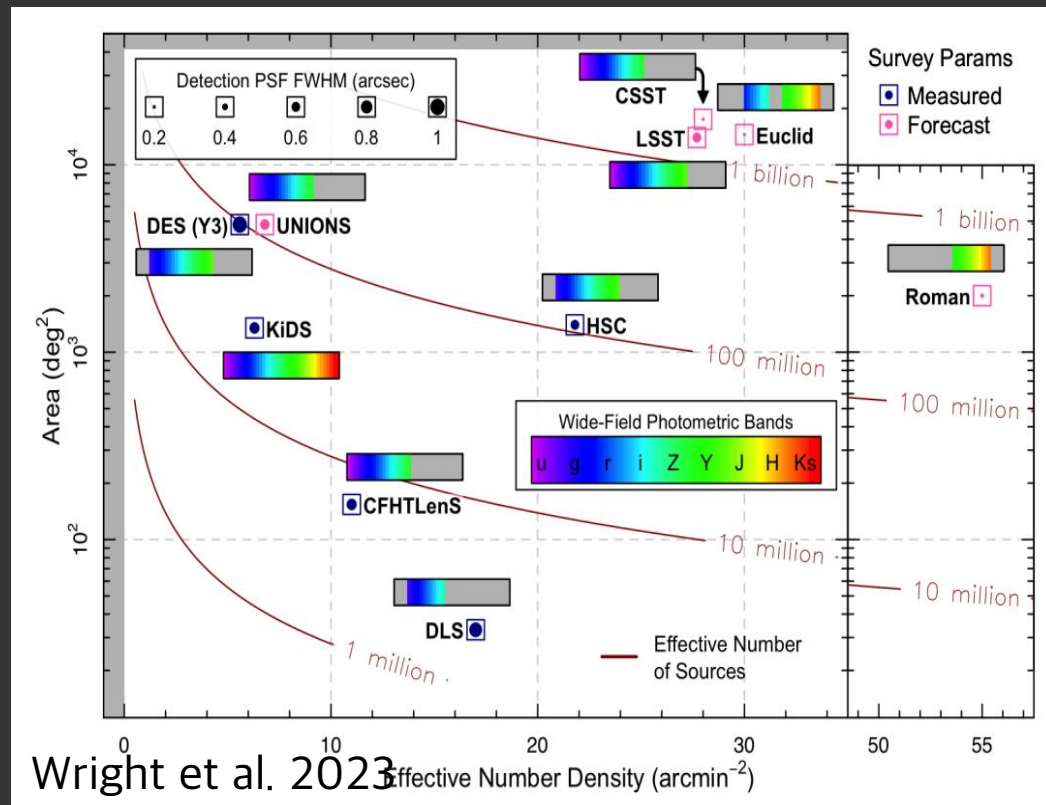


- Chinese Space-station Survey Telescope (CSST)
- 2-Meter Nuv-Optical Telescope
- Launch ~2027

CSST Wide Survey

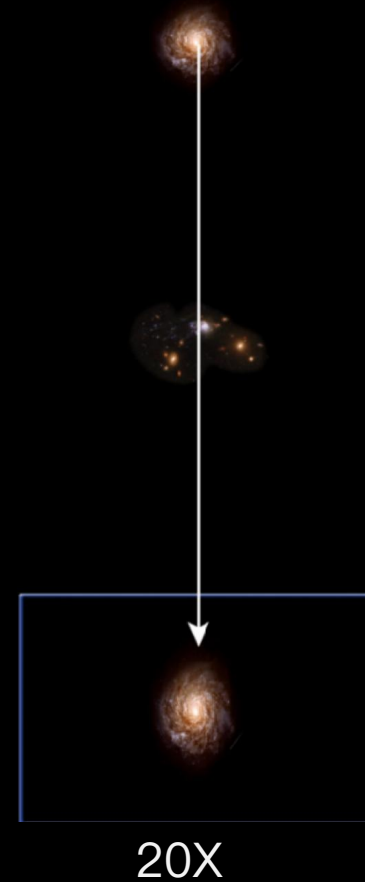
17500 $\square^\circ$ imaging : 255–1000nm, ≥ 6 filters, avg ≥ 25.5 m (5σ , point source, AB mag)

17500 $\square^\circ$ slitless spect: 255–1000nm, 3 bands, $R \geq 200$, ≥ 22 –23m

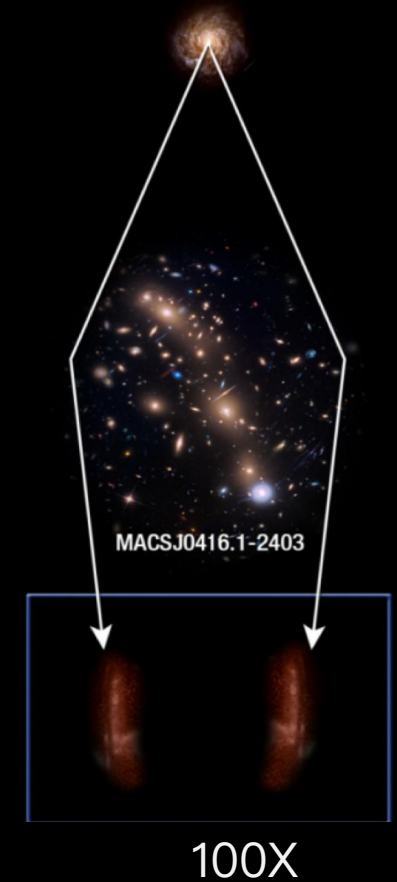


Gravitational Lensing

Weak lens



Strong lens

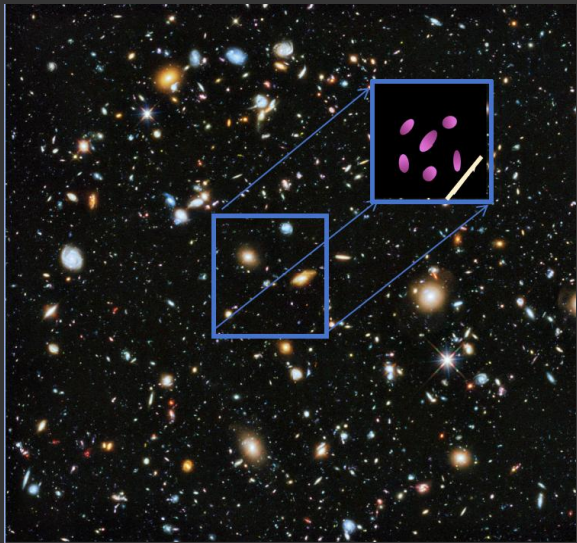


Credit: Nastassia Grimm

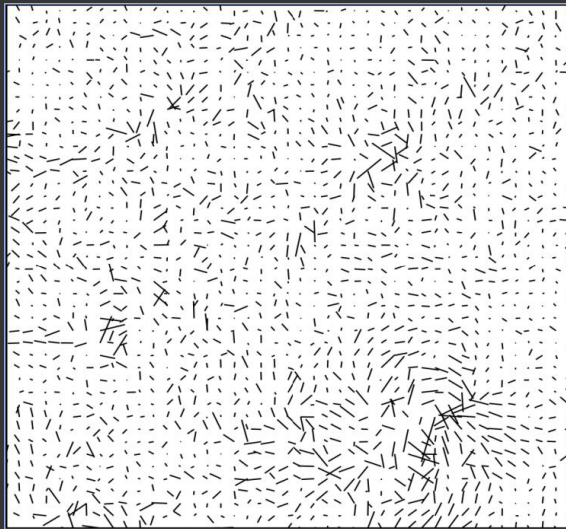
How can AI help?

- Constructing the dark matter map

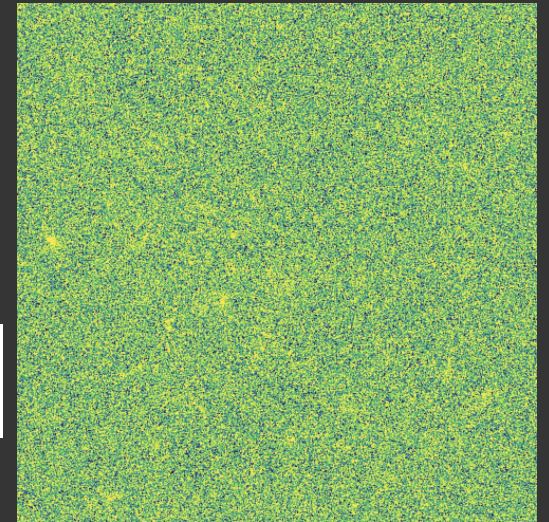
Galaxy shapes



Shear (γ) Field
Derived from galaxy shapes



Convergence k
(Dimensionless DM Surface Density)

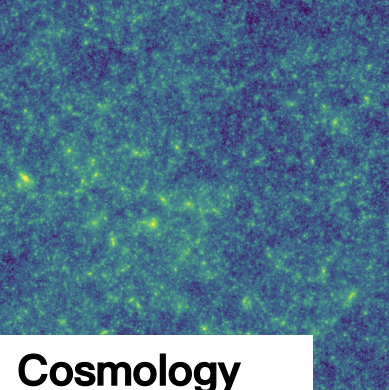


$$\frac{1}{\pi} \int_{\mathbb{R}^2} d^2\theta' \mathcal{D}^*(\theta - \theta') \gamma(\theta')$$

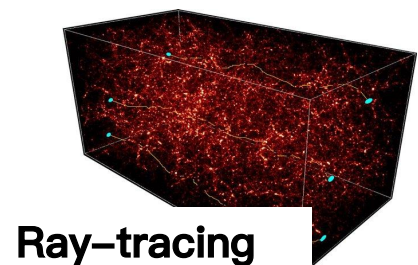
Traditional KS method:

- Noise-dominated on small scales
- Mask-induced artifacts and spurious structures
- Rigid objective: no physics priors, no multi-probe inference

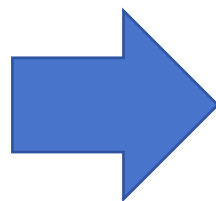
Dark Matter Mass Map with U2Net



Cosmology
Simulation

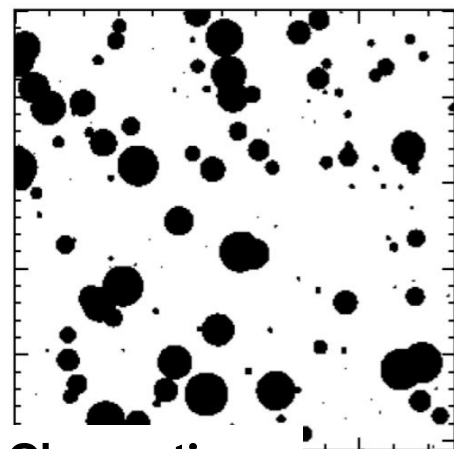
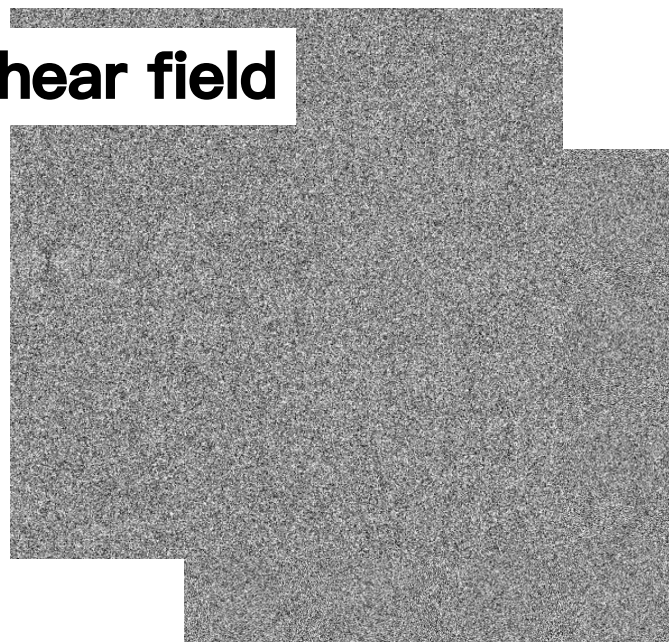


Ray-tracing



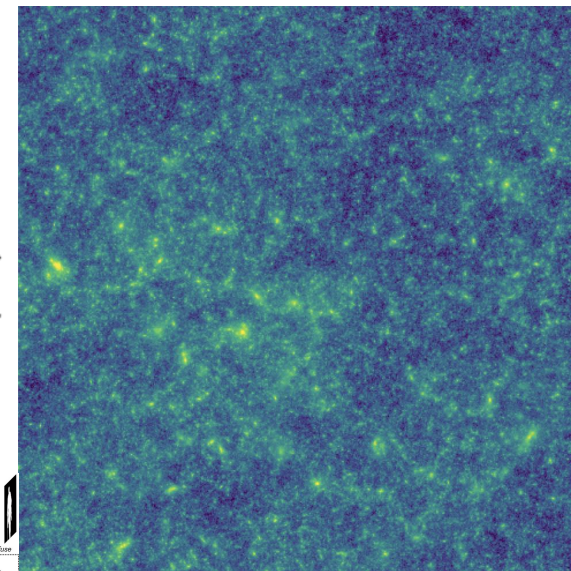
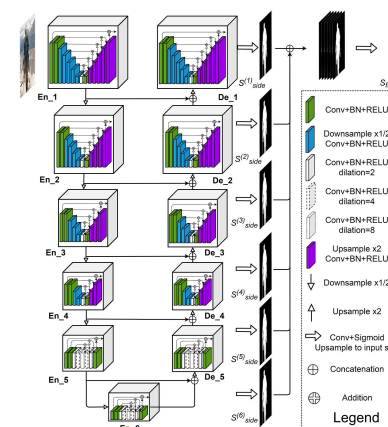
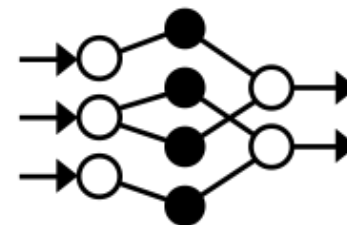
Shear field

x 20000

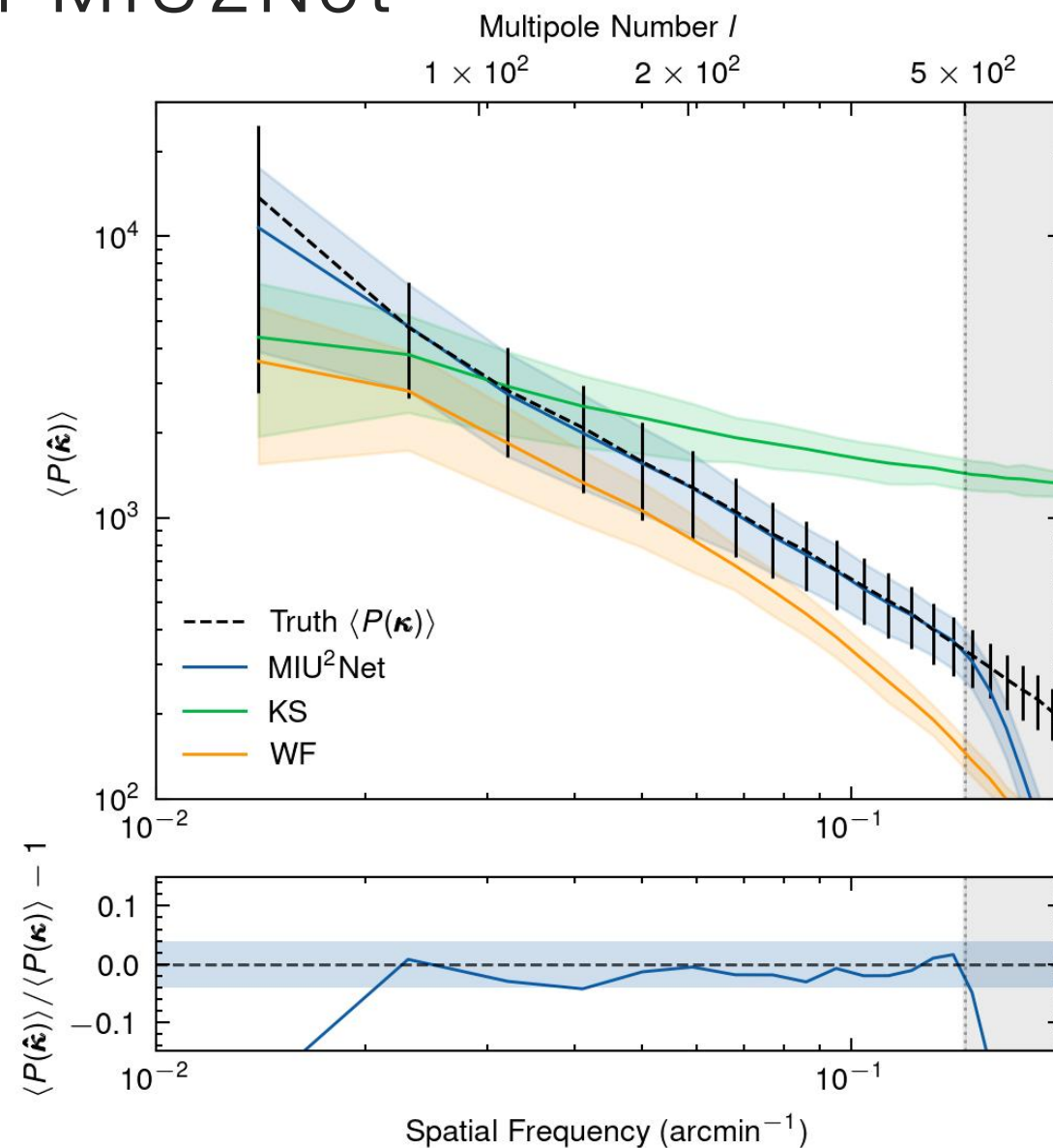
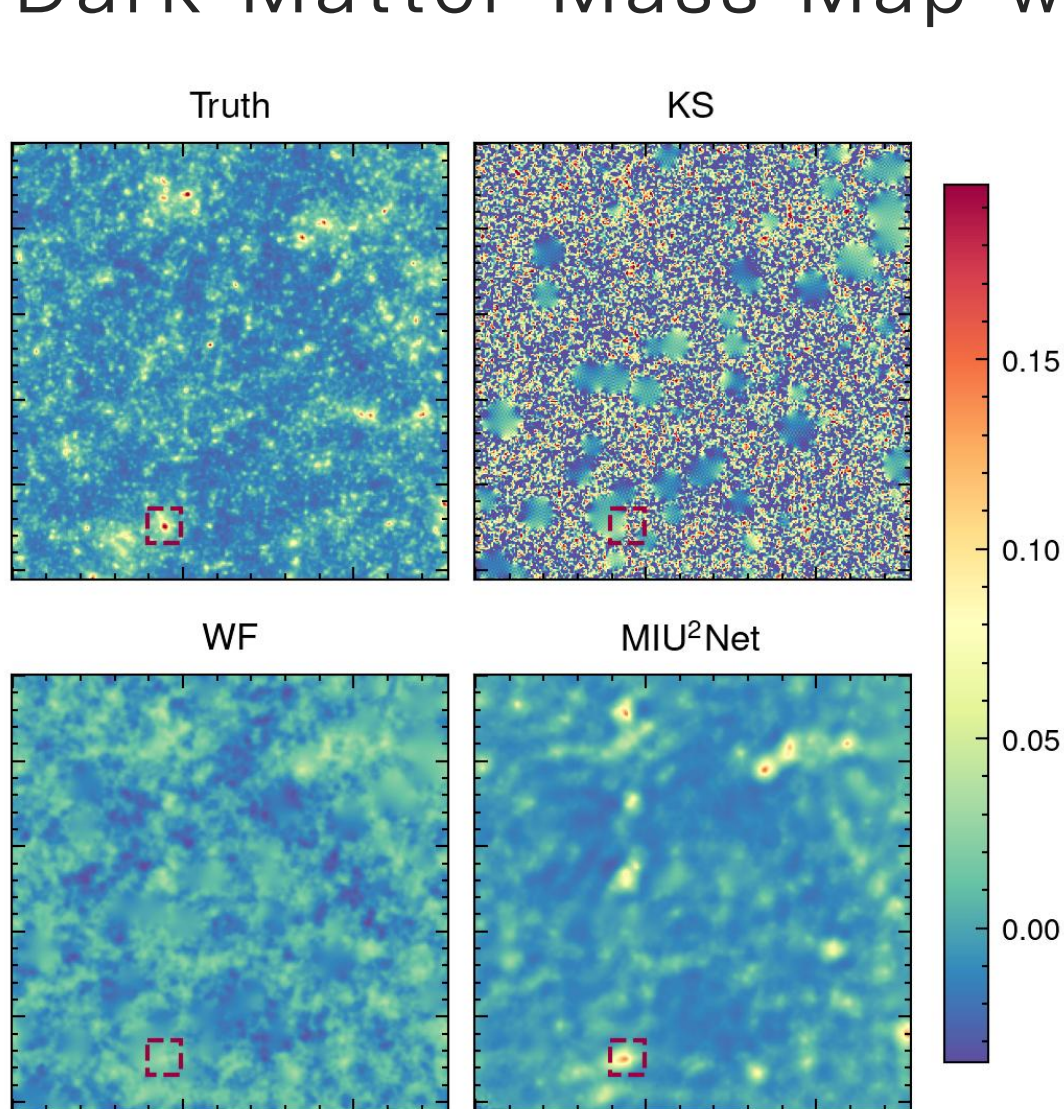


Observation
Mask

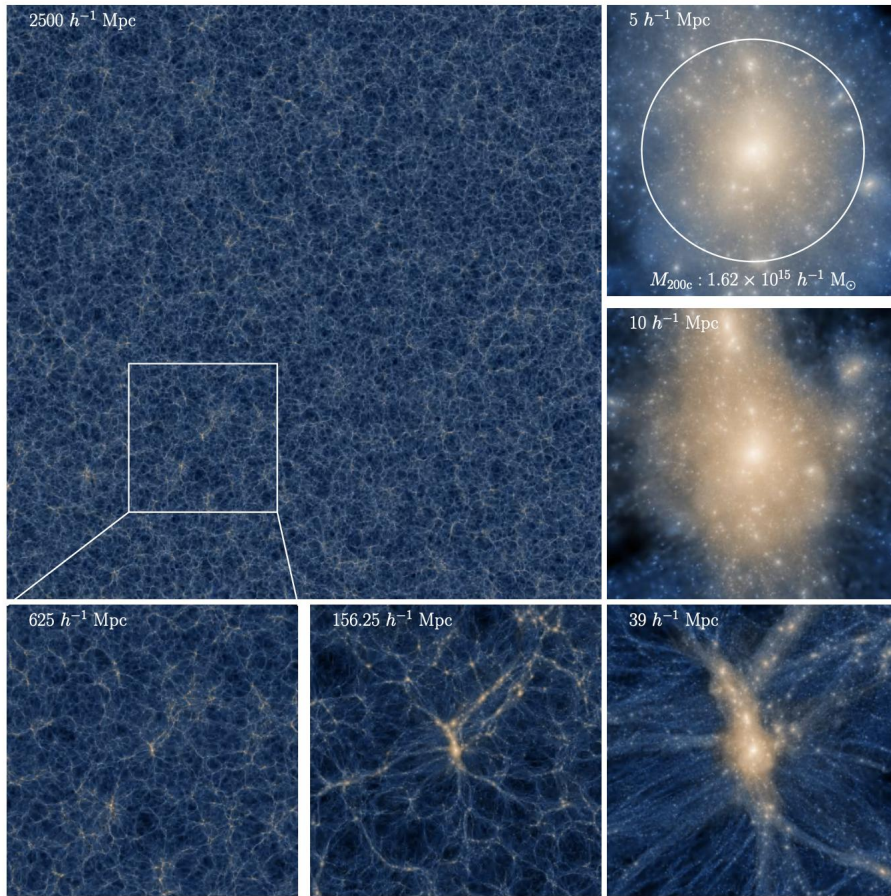
$$\mathcal{H}(\gamma)$$



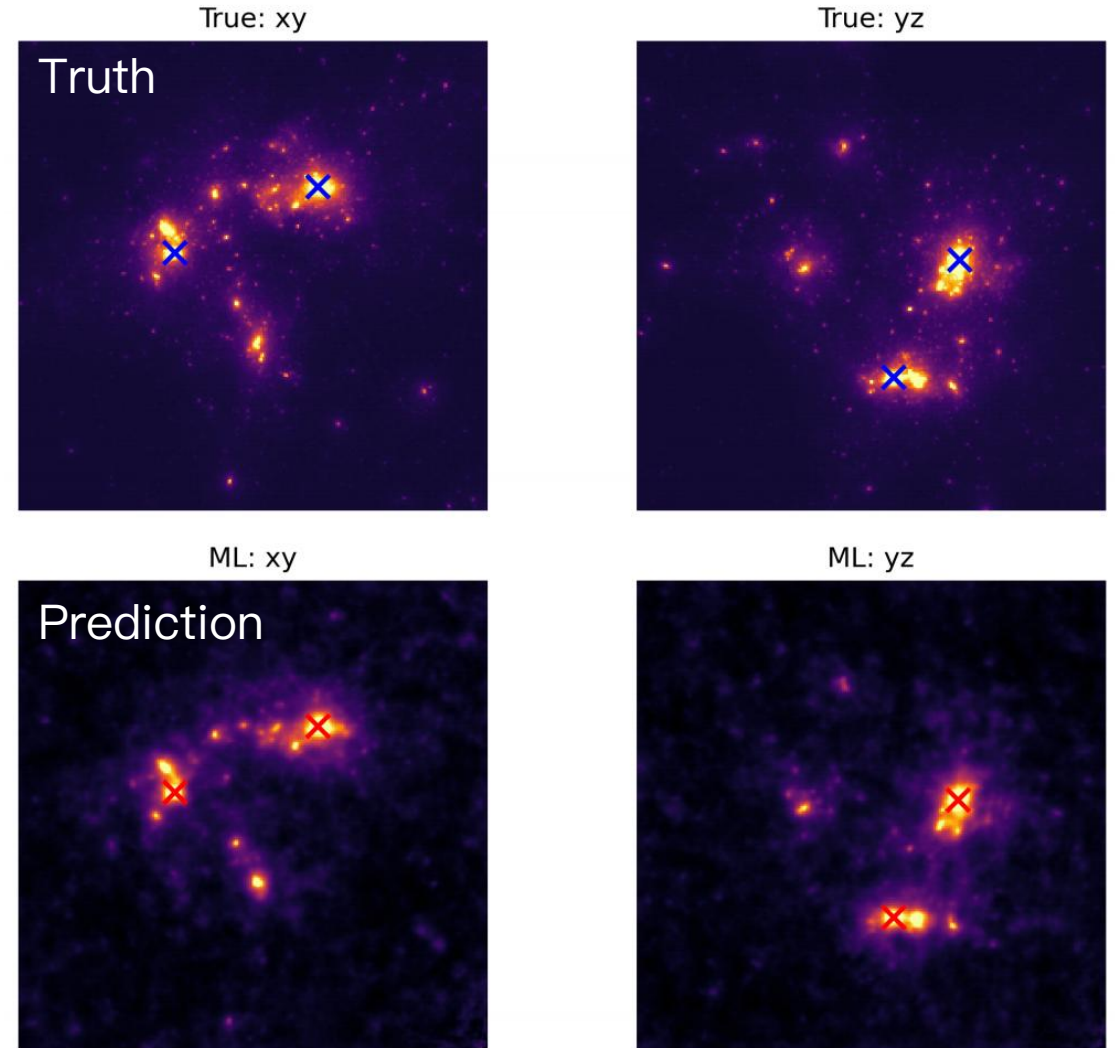
Dark Matter Mass Map with MiU2Net



Cluster-Cluster Pair mass mapping

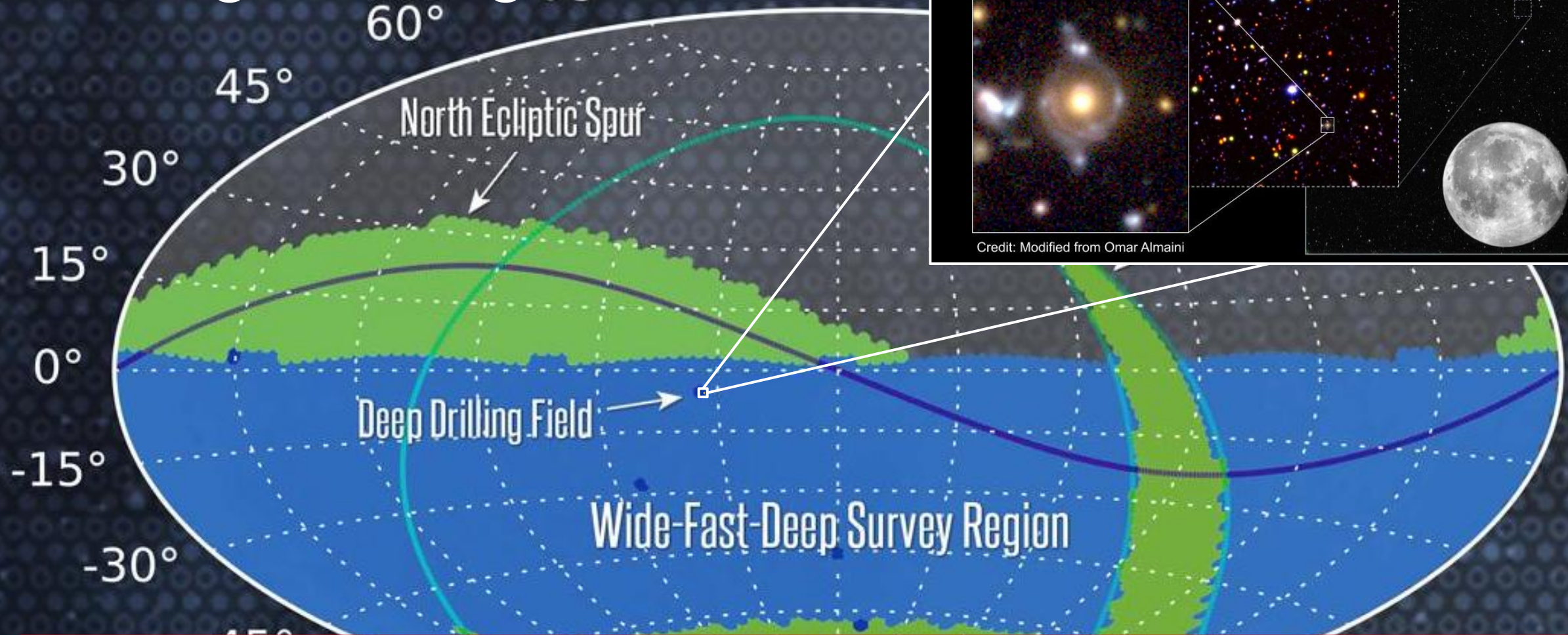


Enhance the training set with HyperMillennium simulation of $(2.5 \text{ Gpc})^3$ box



Chen XY et al. in prep

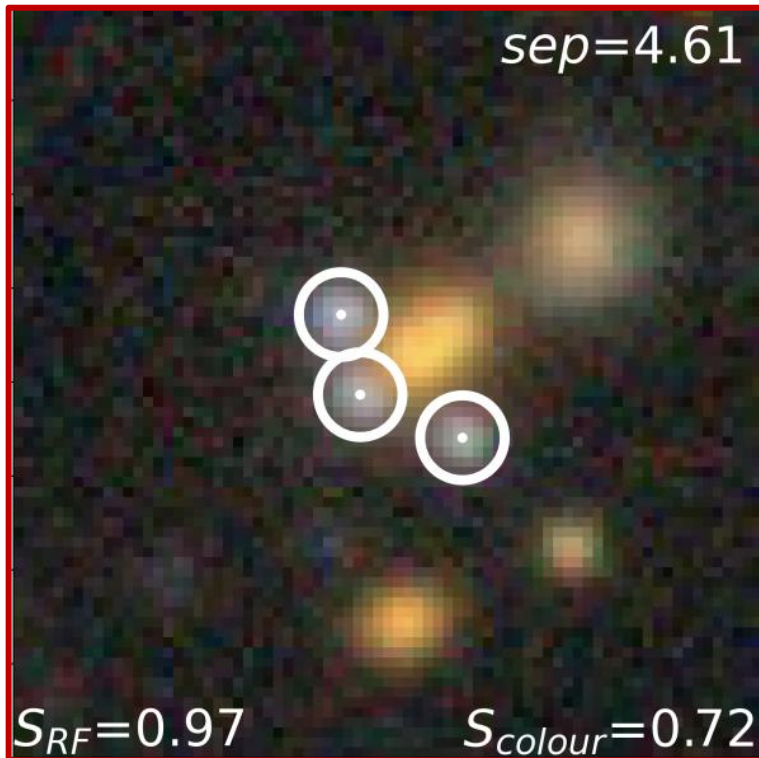
Strong Lensing



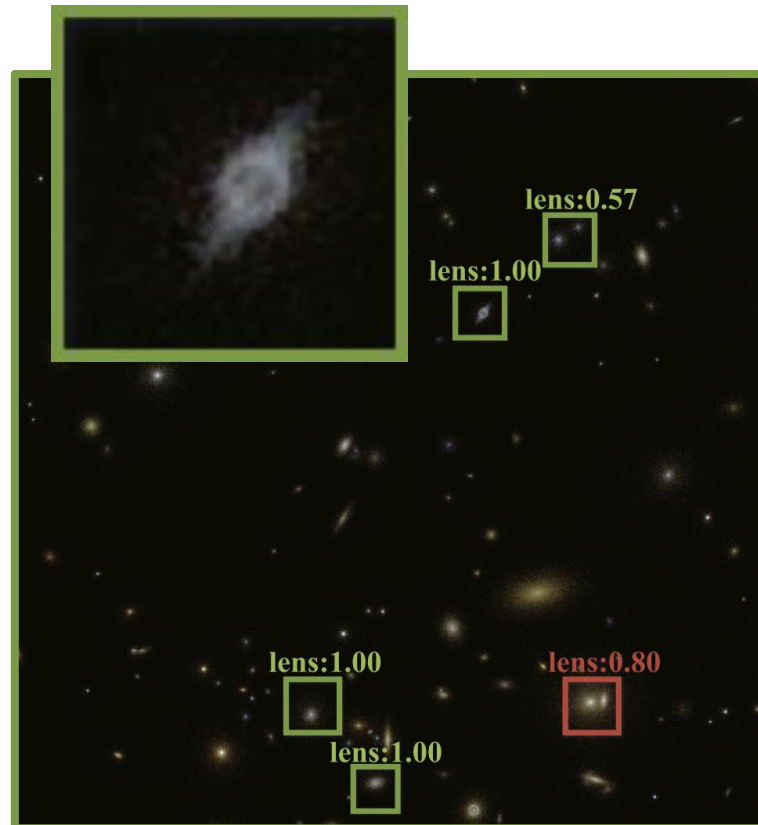
How to identify and model **tens of thousands** of strong lenses from **tens of billions** of astronomical objects in the big data era ???

Identifying Strong Lenses from Gigantic Datasets

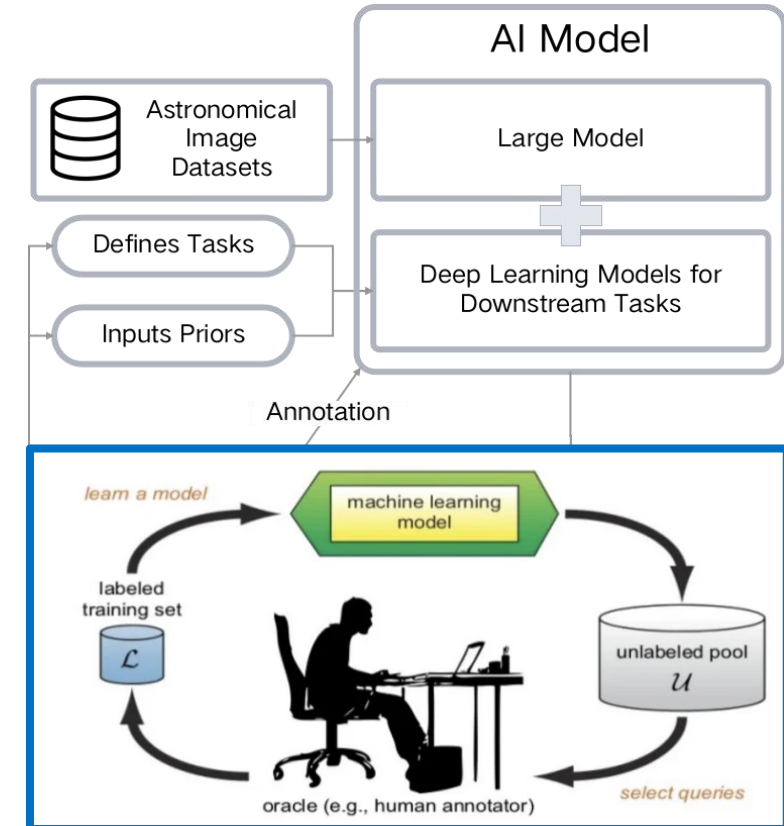
1. Catalog Based Pattern Match



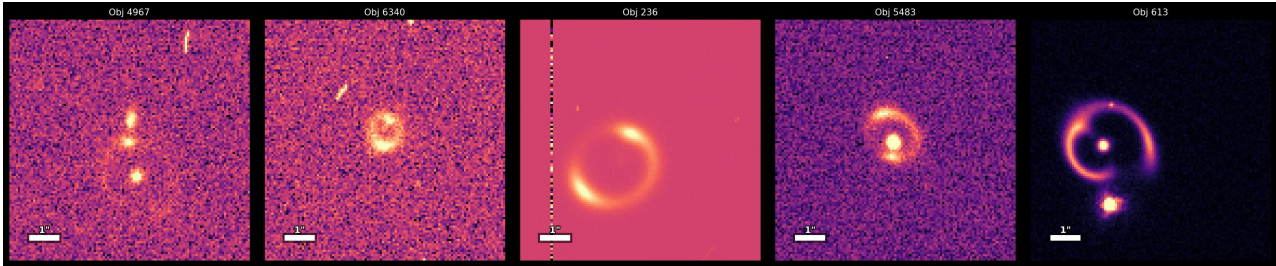
2. Deep Learning Empowered



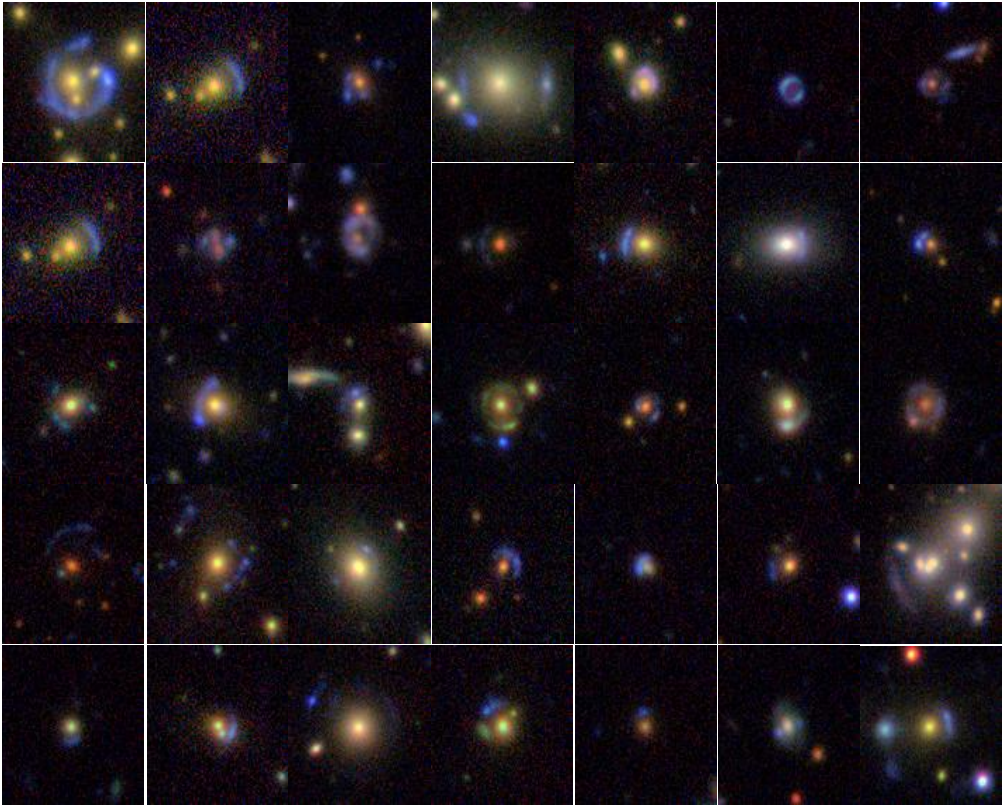
3. Human & AI in the Loop



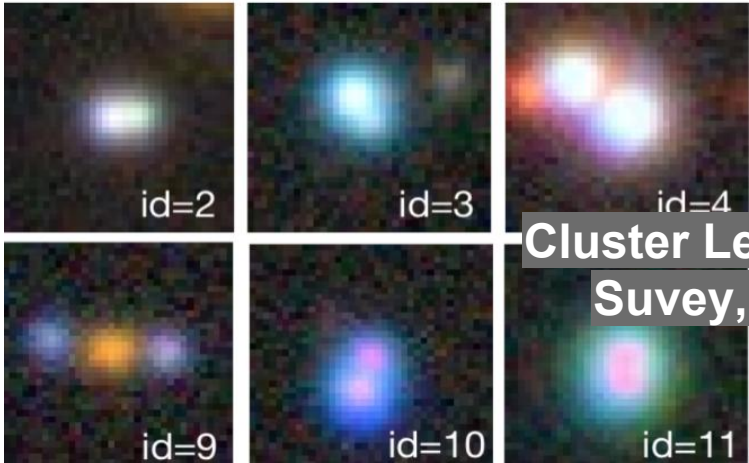
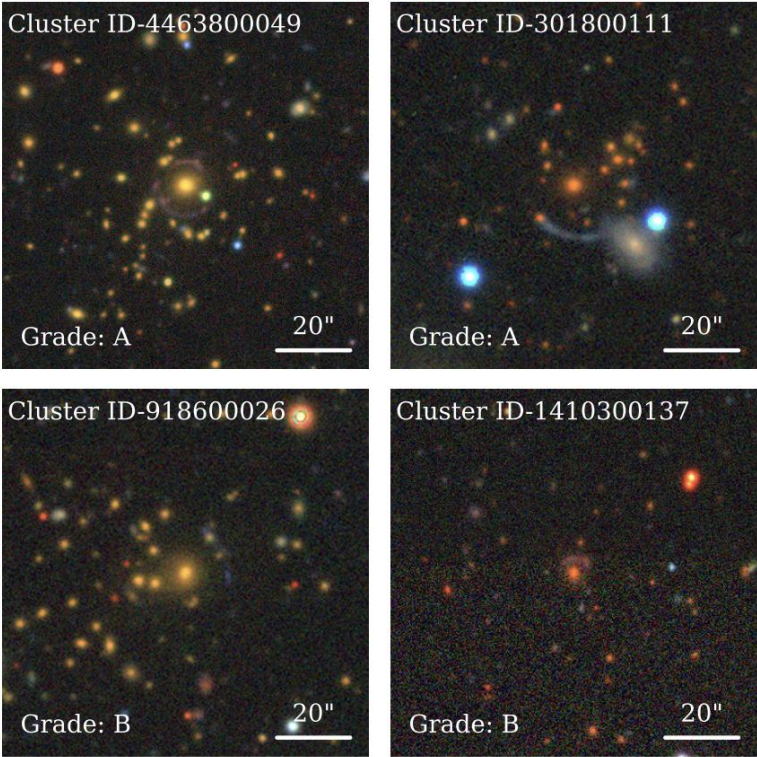
CSST Mock Lenses, Liu P., RuiLi, Cao XY, RL in prep



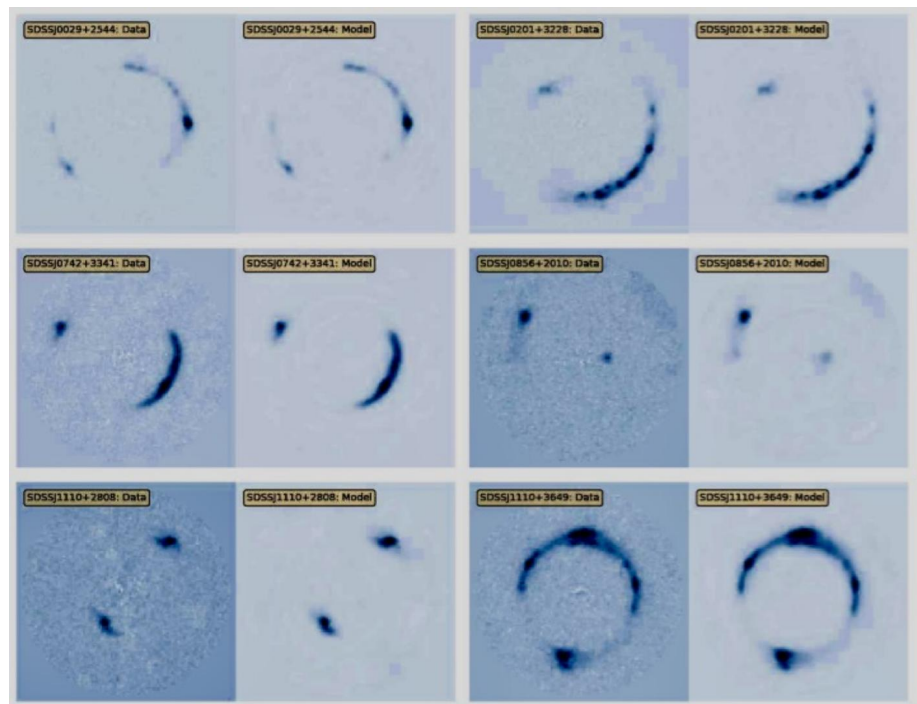
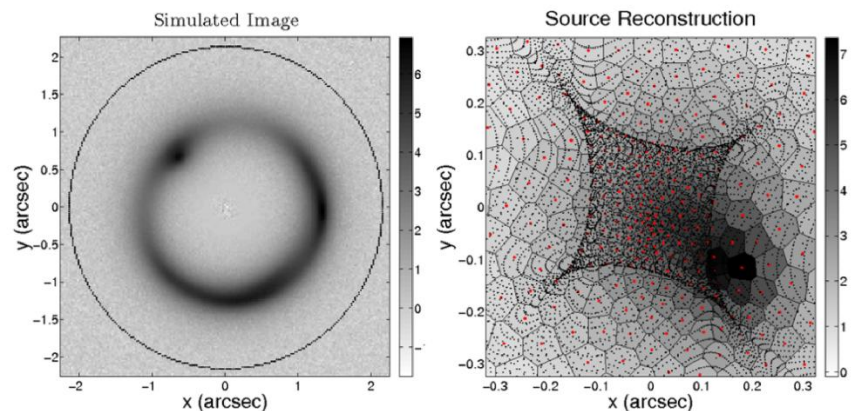
Lens in KiDS DR5(Li Rui et al. 2021)



Cluster Lens in DESI Legacy Suvey, Zhang, NL* et al. 2025

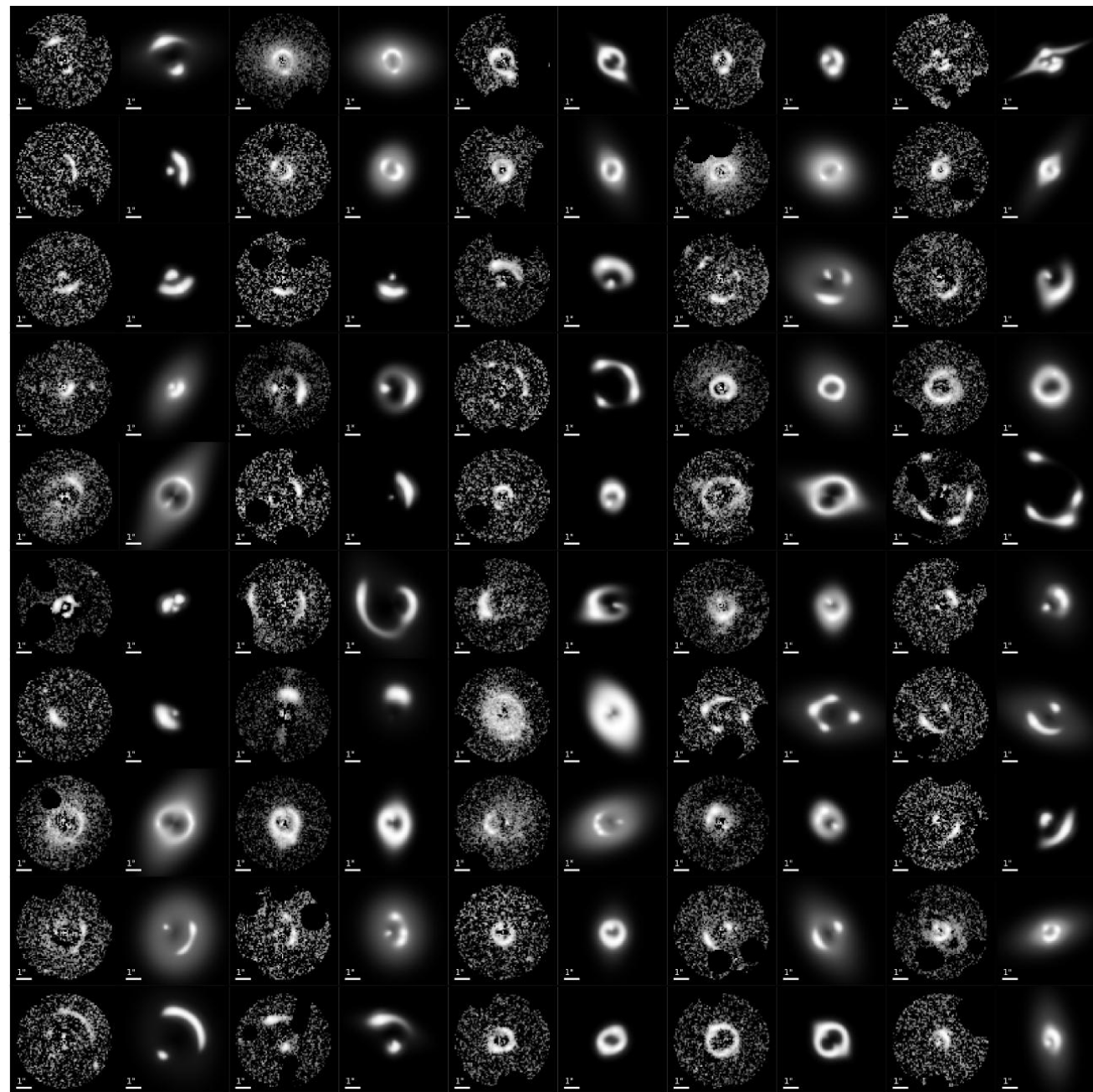


Cluster Lens in DESI Legacy Suvey, He, NL* et al. 2025



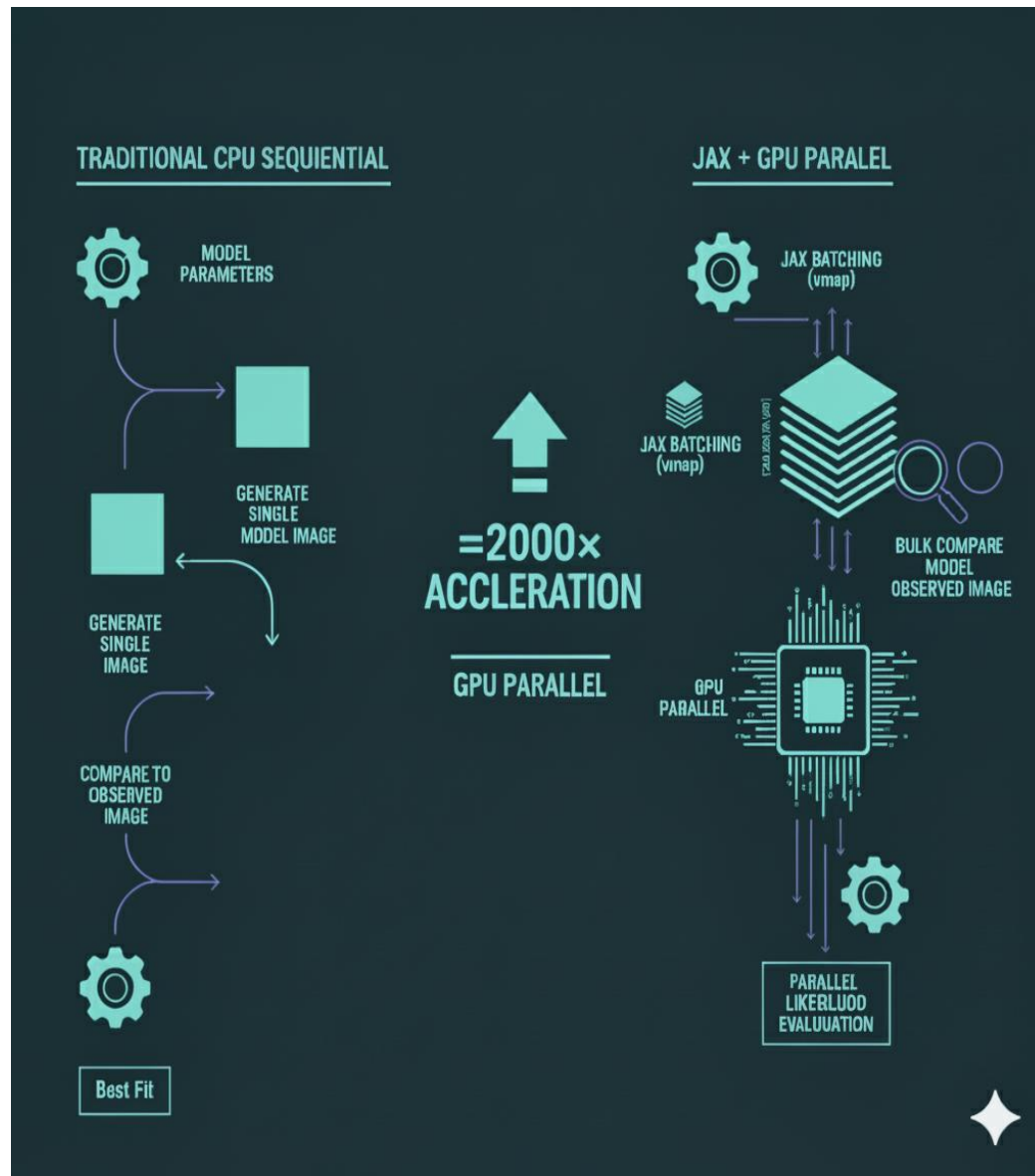
Bells Gallery Lens from Hubble

3/30/2026



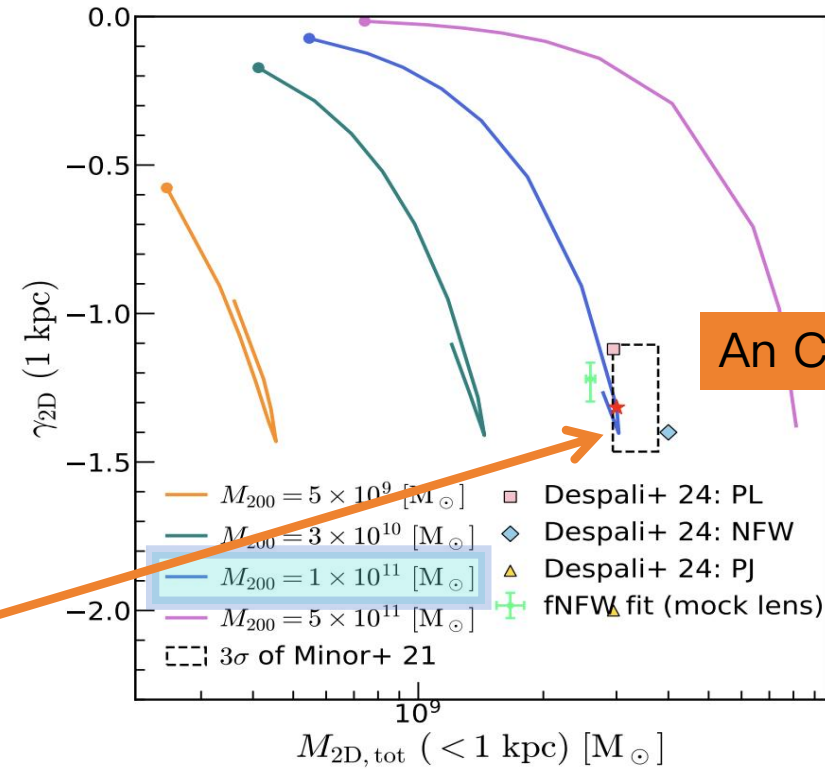
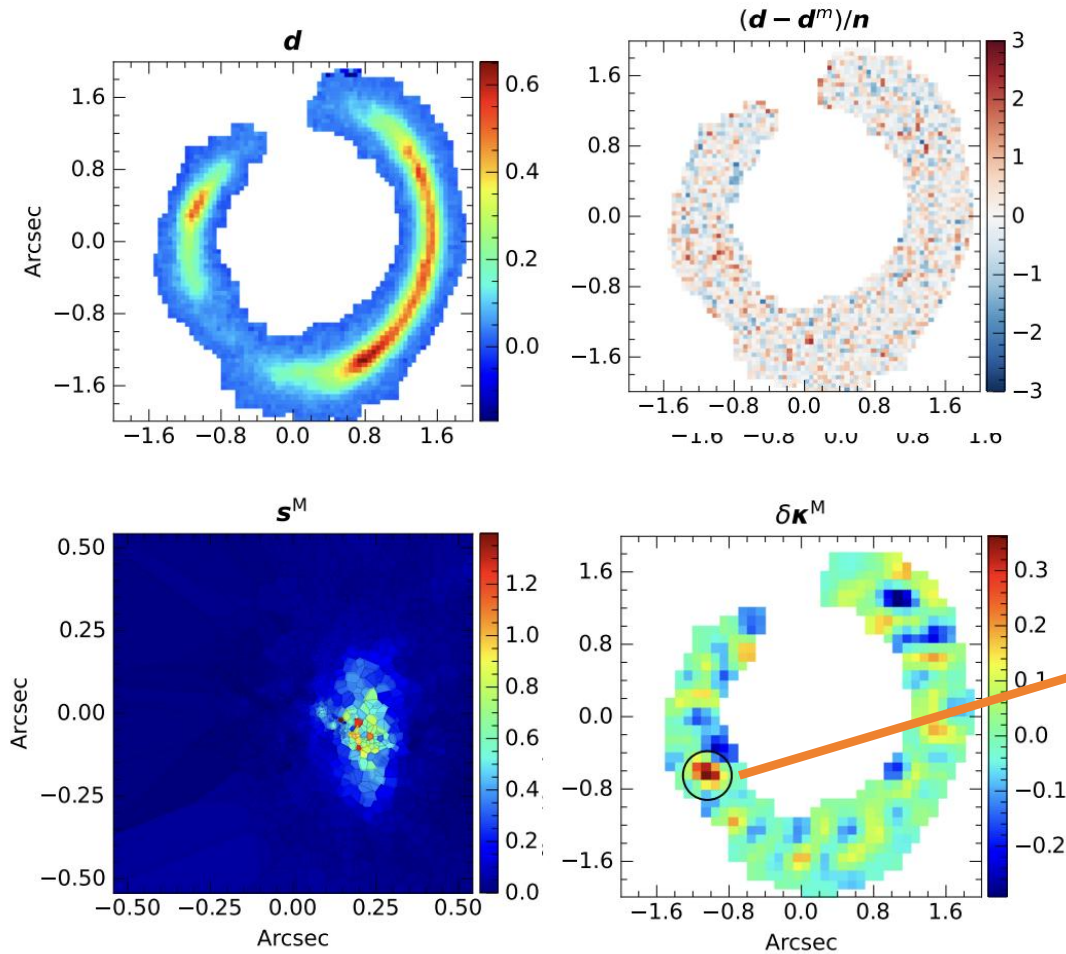
Euclid Q1 lenses model by TINYLENS

16



Reveal the low mass dark matter perturber

SLACS 0946+1006 with Hubble Imaging



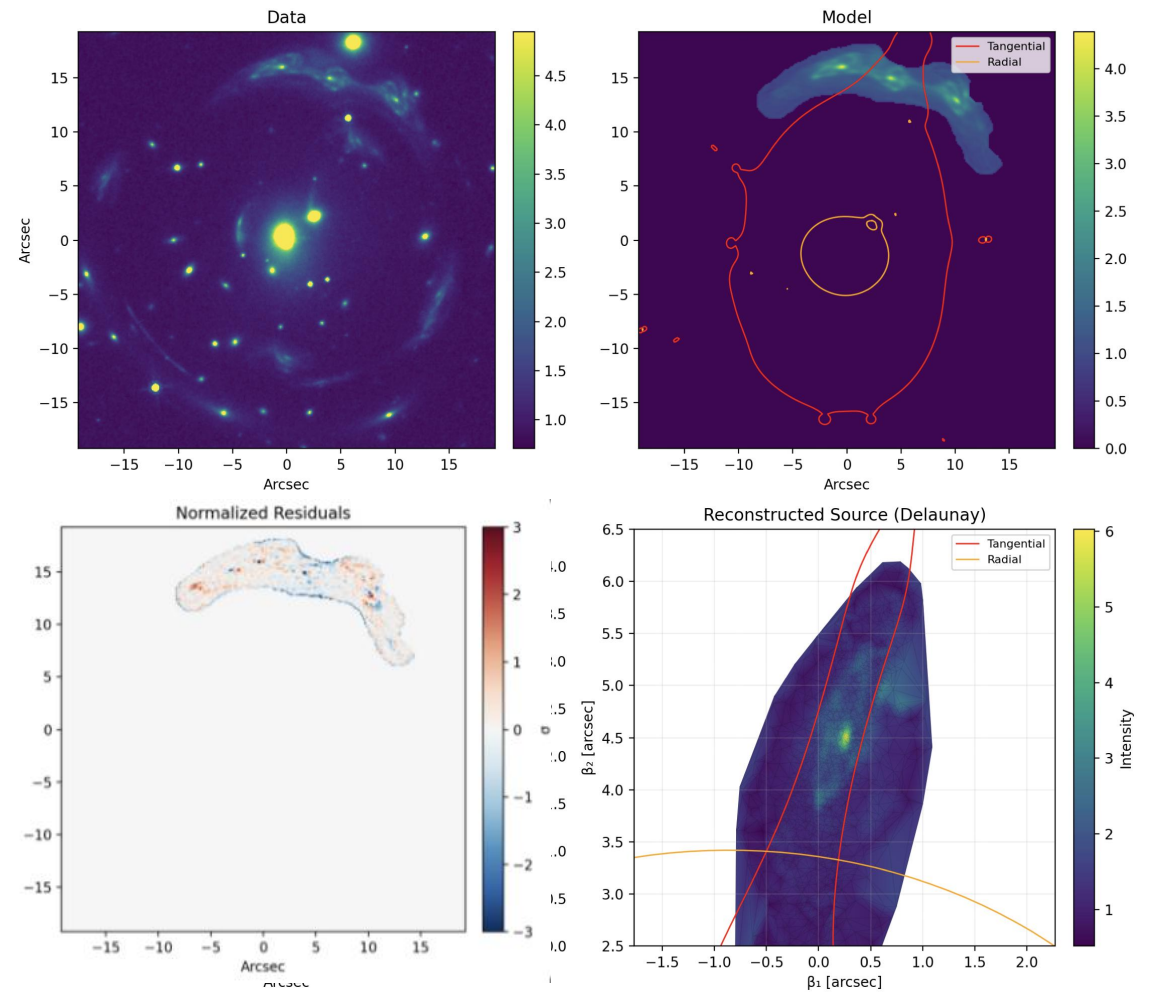
An Challenge for CDM ?

SHB Li, Ran Li + 2025, ApJ, arxiv:
2504.11800

Cluster Giant Arc Modelling



Carousel cluster



Ma XH, RL et al. in prep

AI Language Model for Coding Assistance

How to fit a strong lens model?



Accelerated Learning

- Code Suggestions
- Debugging Help



Before: 1-2 Years of Technical Training



OpenCode Open Source Coding Agent



LensMod: Strong Lens Modelling Project Plan

< >



Discuss Project Requirements



Let's Discuss Project Requirements

Let's discuss the requirements for your strong lens modelling project...



Develop Project Plan



• Great! Here's a plan for your project

It included 5 steps your project. includes development tested.



Define API Interfaces



• I will generate the API interfaces

It included the API interfaces needed for your project.



Assine with Codings



• Assist with Coding

I've written the first part of the code for fitting the strong lens couer.



Assist Using Coding



• Help Design & Test

It's let's set up set testing procedures co. improve performance!



- Quick Project Setup
- Ready to Use Notebooks



Now: 1-2 Months of Learning

AI CAN DO

- High-fidelity dark matter reconstruction
- “Needle in a haystack” into scalable discovery
- Democratizes high-performance computing
- AI Agents redefine how science is done

AI CANNOT DO

- AI does not own reality
 - Reality is not a dataset — it is a messy, constrained system
- AI alone cannot guarantee precision at the level cosmology demands
- AI can not define what is scientifically interesting

In the age of AI, the bottleneck is no longer computation
— it is curiosity



