

# a de Sitter anti-scrambling algebra

David Kolchmeyer

IAS

Based on W.I.P. with Wentao Cui

Observers, wormholes and complex saddles in cosmology

EPFL

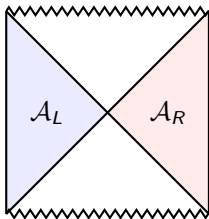
May 19, 2026

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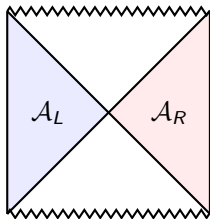
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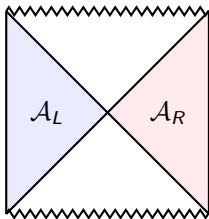
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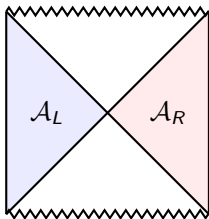
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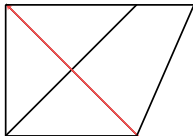
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- ▶ A subalgebra gives a precise notion of a subsystem (or horizon) in quantum gravity, even when quantum metric fluctuations and backreaction are taken into account.

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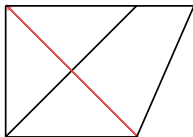
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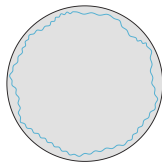


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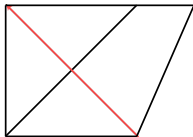


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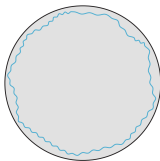


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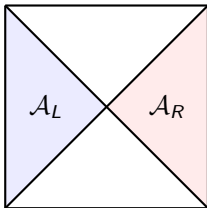


- ▶ By studying algebras in AdS/CFT, we can motivate the duality from the bottom-up.

# Introduction and Summary

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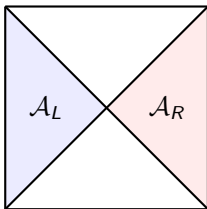
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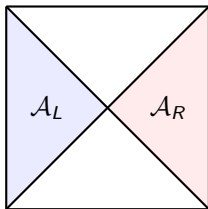


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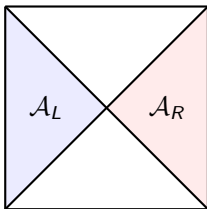


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- ▶ Taken as evidence for static patch holography.
- ▶ I will study  $G_N$  corrections to this picture, following the methods of [Penington-Tabor '25]. We work in a strict  $G_N \rightarrow 0$  limit, holding  $x := G_N e^{\frac{2\pi T}{\beta c}}$  fixed. Operators can have time separations of order  $T$ . We can work exactly in  $x$ , or order by order in  $x$ .

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- ▶ In QFT in dS, the Hartle-Hawking vacuum state is thermal (or KMS) with respect to time translations along the observer's worldline. This is a consequence of the temperature of the geometric horizon. Working exactly in  $x$ , the time-translation symmetry persists, but the KMS condition fails. Order by order in  $x$ , KMS holds.

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- ▶ What does this imply for observer-centric dS holography?

[Anninos-Hartnoll-Hofman '11] [Witten '23] [Narovlansky-Verlinde '23] [Harlow-Usatyuk-Zhao '25]  
[Abdalla-Antonini-Iliesiu-Levine '25] [Akers-Bueller-DeWolfe-Higginbotham-Reinking-Rodriguez '25]  
[Narovlansky '25] [Engelhardt-Gesteau-Harlow '25] [Harlow '26] [Zhao '26] [Goto-Milekhin-Verlinde-Xu '26]

# An observer in de Sitter space

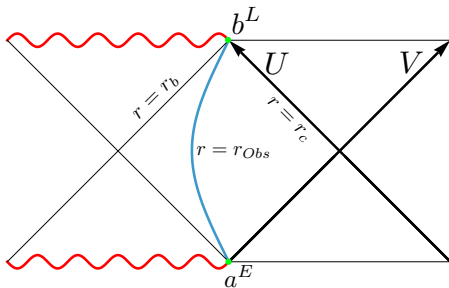
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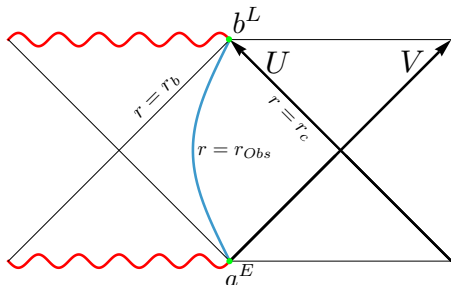
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# An observer in de Sitter space



- ▶ The observer's algebra  $\mathcal{A}_0$  is generated by local QFT operators between  $r_{Obs}$  and  $r_c$ . To compute correlation functions, Wick rotate the time  $t$  to  $-i\tau$ , and compactify  $\tau \sim \tau + \beta_c$ . The Hilbert space  $\mathcal{H}_0$  can be recovered from the GNS construction.

# An observer in de Sitter space

- ▶ We are interested in correlation functions of operators with large time separations. Given  $\mathcal{O} \in \mathcal{A}_0$ , define

$$\mathcal{O}^E := e^{-iHT/2} \mathcal{O} e^{+iHT/2},$$

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- ▶ We will use a doubled Hilbert space

$$\mathcal{H} := \mathcal{H}_A \otimes \mathcal{H}_B,$$

where  $\mathcal{A}_{0,A}$  acts on  $\mathcal{H}_A$  and  $\mathcal{A}_{0,B}$  acts on  $\mathcal{H}_B$ .

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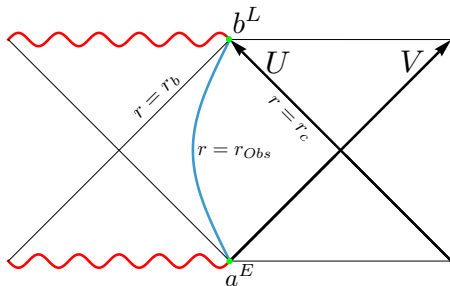
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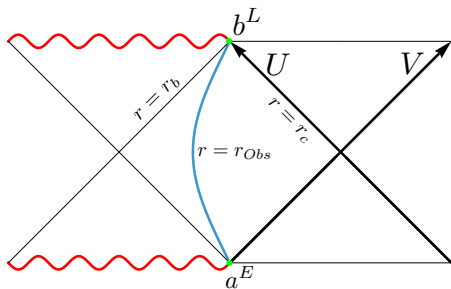
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- Focus on the early operators. We make an ansatz:

$$T_{UU} = -\frac{1}{r_c^2} P_U(\Omega) \delta(U),$$

where  $P_U(\Omega)$ , for every  $\Omega \in S^2$ , is an ANEC operator,

$$P_U(\Omega) := -r_c^2 \int_{-\infty}^{\infty} dU T_{UU}(U, V = 0, \Omega) \sim O\left(e^{\pi T/\beta_c}\right).$$

# An observer in de Sitter space

- ▶  $T_{UU}$  sources a perturbation in linearized gravity with positive c.c.  
( $\Lambda = \frac{3}{\ell_{dS}^2}$ )

$$ds^2 = ds_{SdS}^2 + \delta(U) f_U(\Omega) dU^2,$$
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- ▶ We send  $G_N \rightarrow 0$  with  $G_N e^{2\pi T/\beta_c}$  fixed.
- ▶ The SdS geometry comes in a one-parameter family.  $r_c$  obeys

$$\frac{1}{\sqrt{\Lambda}} < r_c < \sqrt{\frac{3}{\Lambda}},$$

and  $(-\nabla_{S^2}^2 + 1 - r_c^2 \Lambda)$  is invertible strictly within this range.

# An observer in de Sitter space

- ▶ With both early- and late-time shocks, the metric is

$$ds^2 = ds_{SdS}^2 + \delta(U) f_U(\Omega) dU^2 + \delta(V) f_V(\Omega) dV^2,$$

and the action is

$$I = 16\pi G \int d\Omega P_V(\Omega) \frac{1}{-\nabla_{S^2}^2 + 1 - \Lambda r_c^2} P_U(\Omega) \sim O(1).$$

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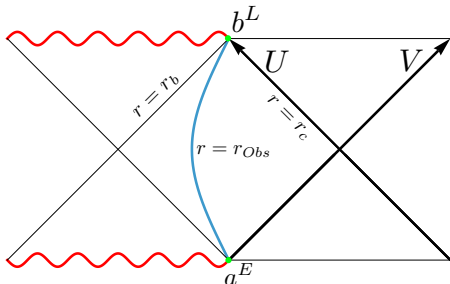
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$$\langle \dots a_1^E b_2^L a_3^E b_4^L \rangle_{\text{QG}} = \langle \dots S a_1 S^{-1} b_2 S a_3 S^{-1} b_4 \rangle_{\text{QFT}}.$$

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Then we may simply write

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- ▶ The observer's algebra  $\mathcal{A}$  is generated by  $\tilde{a}$  and  $\tilde{b}$  for all  $a \in \mathcal{A}_{0,A}$  and all  $b \in \mathcal{A}_{0,B}$ .

# An observer in de Sitter space

- ▶ Recall that

$$\frac{1}{\sqrt{\Lambda}} < r_c < \sqrt{\frac{3}{\Lambda}}.$$

- ▶ Consider the limit  $r_c \rightarrow \sqrt{\frac{3}{\Lambda}}$  with  $G_N \rightarrow 0$  and  $M$  fixed. The SdS geometry becomes global dS. The observer sits at  $r = 0$ . The eikonal phase becomes

$$\delta = \frac{2\ell_{dS}}{M} e^{2\pi T/\beta_c} \int d\Omega_1 d\Omega_2 P_U(\Omega_1) P_V(\Omega_2) \cos \Theta(\Omega_1, \Omega_2).$$

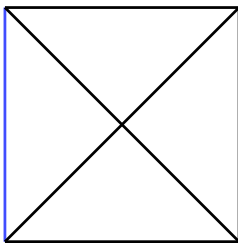
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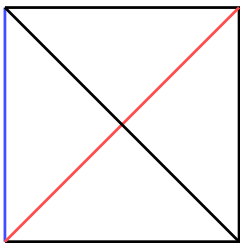
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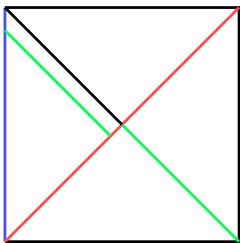
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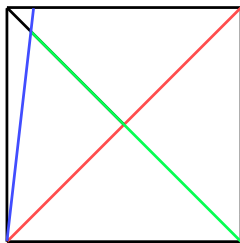
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- ▶ The  $s$ -wave sector decouples.

$$\delta = -\frac{4\pi}{\Delta S_c} \frac{e^{2\pi T/\beta_c}}{\Lambda} P_U P_V,$$

with

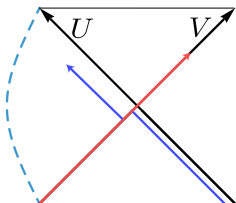
$$P_U := \int d\Omega P_U(\Omega), \quad P_V := \int d\Omega P_V(\Omega)$$

# A toy model of the s-wave sector

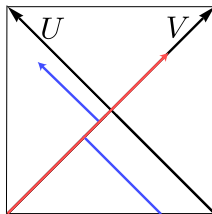
- As a toy model of the s-wave sector, we will take QFT A to be a right-moving 2D chiral CFT. QFT B will be a left-moving 2D chiral CFT. The S matrix is

$$S := e^{ixP_A P_B},$$

where  $x$  is a parameter, proportional to  $\pm G_N e^{\frac{2\pi T}{\beta_c}}$ .



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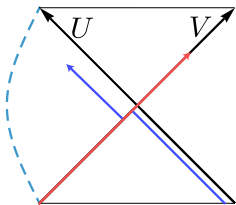
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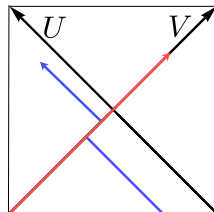
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We will study this model both exactly in  $x$ , and perturbatively in  $x$ .

- ▶ Fields are functions of the affine coordinate on the horizon,  $U$ . A primary field  $\phi_j(U)$  with weight  $h_j \in \mathbb{Z}_{\geq 0}$  transforms as

$$e^{iyP} \phi_j(U) e^{-iyP} = \phi_j(U - y)$$

$$e^{iyB} \phi_j(U) e^{-iyB} = e^{-h_j y} \phi_j(e^{-y} U)$$

$$e^{iyK} \phi_j(U) e^{-iyK} = \left( \frac{1}{1 - Uy} \right)^{2h_j} \phi\left( \frac{U}{1 - Uy} \right)$$

The Hermitian generators  $P$ ,  $B$ , and  $K$  obey

$$[B, P] = iP,$$

$$[B, K] = -iK,$$

$$[P, K] = 2iB.$$

The primaries are Hermitian w.l.o.g.

I will suppress the subscripts on the fields when it will not cause confusion.

- ▶ Spacelike separated operators commute,

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- ▶ The Hilbert space is generated by the vacuum,  $|\Omega\rangle$ , and  $\phi_i(U) |\Omega\rangle$  for  $U$  in any interval in  $\mathbb{R}$  (Reeh–Schlieder theorem and state-operator correspondence).

# The commutant of the observer's algebra

- ▶ For  $x > 0$ , the operators  $\tilde{a}$  and  $\tilde{b}$  for  $a \in \mathcal{A}_{0,A}$  and  $b \in \mathcal{A}_{0,B}$  can be interpreted as operators dual to insertions on an AdS boundary.

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- ▶ For  $x < 0$  (dS), we may say that  $\tilde{a}'$  and  $\tilde{b}'$  generate  $\mathcal{A}'$  order by order in perturbation theory. There is a pair of factorizing algebras.

# The KMS condition in 2D Chiral CFT

- ▶ Consider a thermal two-point function in finite-dimensional quantum mechanics:

$$F(t_1, t_2) = \text{Tr} e^{-2\pi H} \mathcal{O}_1(t_1) \mathcal{O}_2(t_2),$$

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$$\langle \phi_1(U_1) \cdots \phi_n(U_n) \rangle = \langle \phi_1(0) e^{iU_1 P} \cdots \phi_n(U_n) \rangle$$

is holomorphic for  $U_1$  in the lower-half-plane, because  $P \leq 0$ . Thus, we may continue  $U_1 \rightarrow U_1 e^{-\pi i} = -U_1$ .

$$\langle \phi_1(U_1) \cdots \phi_n(U_n) \rangle \rightarrow \langle \phi_1(e^{-i\pi} U_1) \cdots \phi_n(U_n) \rangle = \langle \cdots \phi_n(U_n) \phi_1(e^{-i\pi} U_1) \rangle$$

Hence, we may analytically continue  $U_1 \rightarrow U_1 e^{-2\pi i}$ , and the KMS condition is established.

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- ▶ The KMS condition involves analytic continuation:  $t \rightarrow t - \beta i$ . Have we analytically continued into a “fake” region? [Lin-Stanford '23].

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- ▶ The KMS condition is tied to Tomita-Takesaki theory, which I will briefly review.
- ▶ A state  $|\Psi\rangle$  is *cyclic* with respect to a von Neumann algebra  $\mathcal{A}$  if the set

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is dense in the Hilbert space.

- ▶ A state  $|\Psi\rangle$  is *separating* with respect to a von Neumann algebra  $\mathcal{A}$  if

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- ▶ **Theorem.** (Reeh-Schlieder): In chiral CFT, the vacuum  $|\Omega\rangle$  is cyclic and separating for the algebra of operators of an interval on the horizon,  $U \in [U_1, U_2]$ .

# Tomita-Takesaki theory in chiral CFT

- ▶ Let  $|\Psi\rangle$  be cyclic and separating for  $\mathcal{A}$ . The *Tomita operator*  $S_\Psi$  is defined by

$$S_\Psi a |\Psi\rangle = a^\dagger |\Psi\rangle.$$

- ▶ The *modular operator* is defined by  $\Delta_\Psi := S_\Psi^\dagger S_\Psi$ .
- ▶ Matrix elements of  $\Delta_\Psi$  are computed using

$$\langle \Psi | a_1 \Delta_\Psi a_2 | \Psi \rangle = \langle \Psi | a_2 a_1 | \Psi \rangle, \quad a_1, a_2 \in \mathcal{A}.$$

- ▶ For a chiral CFT, let  $\mathcal{A}$  be the algebra of operators associated to  $U > 0$ . Then  $\Delta_\Omega = e^{2\pi B}$ . Time translation is given by  $a(t) = e^{-itB} a e^{itB}$ . Thus,

$$\langle \Psi | a_1(-2\pi i) a_2 | \Psi \rangle = \langle \Psi | a_2 a_1 | \Psi \rangle,$$

which is the KMS condition.

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- ▶ **Conjecture.**  $\mathcal{A}' = \mathbb{C}$ .

# Summary and Conclusion

- ▶ A static observer at the North Pole in dS has a horizon, and an associated static patch algebra  $\mathcal{A}$ . The commutant  $\mathcal{A}'$  represents the other static patch.  $\mathcal{A}$  and  $\mathcal{A}'$  are a pair of factorizing algebras, and vacuum correlators obey the KMS condition.

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