

Dilaton gravity in box: canonical quantization of JT and sinh models

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Observers, wormholes and complex saddles in cosmology,
EPFL, 20 May 2026

WIP with L. Russo (Florence), A. Tarana (Parma), L. Griguolo (Parma),
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Introduction and motivation

Finite cutoff quantum gravity: define **bulk observables** associated with boundary observers located at finite radial location

Combine with low dimensional (2d) gravity framework

- **Most renowned example: JT gravity**

→ duality on the disk with boundary reparametrization mode captured by **Schwarzian theory** → low-energy limit of SYK model

→ what happens if we push the boundary clock into the bulk?

- 2 other solvable theories: **Liouville and complex Liouville gravity**

→ non-critical string theories given by **double 2d Liouville CFT**

→ Dilaton gravity perspective: through field redefinition

$$\phi = b^{-1}\rho - b\pi\Phi, \quad \chi = b^{-1}\rho + b\pi\Phi,$$

mapped to 2d dilaton gravities with **sinh/sine potential**

→ **asymptotic dualities** with **DSSYK/ q -Schwarzian deformed?**

Plan of the talk

JT gravity

- Classical JT gravity as a scattering problem at finite cutoff
- Canonical quantization and partition function
- Embedding into $SL(2, \mathbb{R})$ group theory
- Exact and semiclassical two-point function

Sinh dilaton gravity (results only)

- Effective geometry Hamiltonian formulation
- Canonical quantization and partition function
- Quantum group $SL_q(2, \mathbb{R})$ perspective

Outlook

JT gravity on shell

$$S_{\text{JT}}[g, \Phi] = \frac{1}{2} \left[\int_M d^2x \sqrt{-g} \Phi (R + 2) + 2 \int_{\partial M} du \sqrt{|h|} \Phi (K - 1) \right]$$

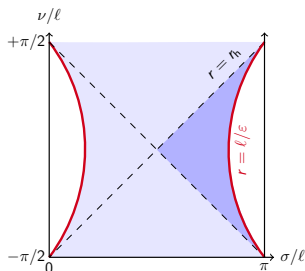
Classical solutions in the Schwarzschild gauge

$$ds^2 = (r_h^2 - r^2) dt^2 - \frac{1}{r_h^2 - r^2} dr^2, \quad \Phi = r$$

Standard formulation: $\epsilon \rightarrow 0$

$$\Phi \Big|_{\partial M} = \Phi_b \sim 1/\epsilon, \quad g_{\mu\nu} dx^\mu dx^\nu \Big|_{\partial M} = \text{const.}$$

- **Goal:** keep Φ_b finite
- Consider the global extension of AdS_2
- **Two-dimensional phase space**



What are the natural **phase space variables** in the two-sided finite cutoff description?

Diff invariant quantity \rightarrow **proper time** $\tau^{L/R}$ of the boundaries

$$\tau^{L/R} \equiv \sqrt{\Phi_b^2 - \Phi_h^2} t^{L/R},$$

Killing vector fields $\zeta_{L/R} \equiv \partial_{\tau^{L/R}} = \frac{1}{\sqrt{\Phi_b^2 - \Phi_h^2}} \partial_{t^{L/R}}$

The corresponding **Hamiltonians** \rightarrow **conserved charges** associated with translations in $\tau^{L/R}$.

$$H^{L/R} = \zeta_{L/R}^\alpha \zeta_{L/R}^\beta T_{\alpha\beta}^{L/R} = \boxed{\Phi_b - \sqrt{\Phi_b^2 - \Phi_h^2}},$$

where $T_{\alpha\beta}^{L/R}$ is the Brown–York stress tensor on each boundary:

$$T_{\alpha\beta}^{L/R} \equiv -\frac{2}{\sqrt{|h|}} \frac{\delta S_{JT}}{\delta h^{\alpha\beta}} = (n^\mu \nabla_\mu \Phi - \Phi) \Big|_{h_{\alpha\beta}}$$

$H = \frac{H^L + H^R}{2}$ is **quasi-local Hamiltonian** conjugate to **Tolman temperature** $\beta_T = \beta$.

Define a finite cutoff **symplectic measure** $\omega = d\tau \wedge dH$ on phase space ¹

¹Harlow, Jafferis.

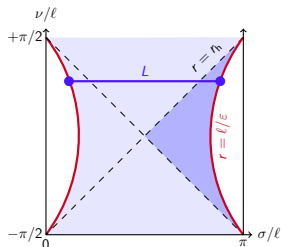
JT gravity as a scattering problem

Natural variable: **geodesic length** of a **two-sided Cauchy slice** at finite cutoff:

$$L = 2 \sinh^{-1} \left[\frac{\sqrt{\Phi_b^2 - \Phi_h^2}}{\Phi_h} \cosh(\Phi_h t) \right]$$

Canonical **change of variables**: $\omega = dL \wedge dP$

$$H(L, P) = \Phi_b - \sqrt{\Phi_b^2 - \left(P^2 + \frac{\Phi_b^2}{\cosh^2(L/2)} \right)}$$



We focus on the **Pöschl–Teller Hamiltonian** inside the square root

$$H_0 = P^2 + \frac{\Phi_b^2}{\cosh^2(L/2)}$$

As $\Phi_b \gg 1$, $L_{\text{ren}} = L - 2 \log(\Phi_b)$, it reduces to standard Liouville Hamiltonian $H_0 \rightarrow \frac{1}{2\Phi_b} (P^2 + e^{L_{\text{ren}}})$.

Canonical quantization

We propose a **quantum ordering** yielding the **Schrödinger equation**:²

$$-\frac{d^2\psi_s(L)}{dL^2} + \left(\frac{\nu^2}{\cosh^2(L/2)} - \frac{1}{16 \sinh^2(L/2) \cosh^2(L/2)} \right) \psi_s(L) = s^2 \psi_s(L).$$

Effective potential with quantum effect $V_{\text{eff}}(L) \stackrel{L \rightarrow 0}{\sim} -\frac{\hbar^2}{4L^2}$.

$$V_{\text{eff}}(L) = \frac{\Phi_b^2}{\cosh^2(L/2)} - \frac{\hbar^2}{16 \sinh^2(L/2) \cosh^2(L/2)},$$

which is precisely induced by $b \rightarrow 0$ limit of sinh dilaton gravity!

The wavefunction is

$$\psi_s(L) = \tanh^{\frac{1}{2}}\left(\frac{L}{2}\right) \cosh^{-2is}\left(\frac{L}{2}\right) {}_2F_1\left(\frac{1}{2} + is + i\nu, \frac{1}{2} + is - i\nu, 1; \tanh^2\left(\frac{L}{2}\right)\right)$$

vanishes at $L = 0$ \rightarrow consistent with the interpretation of L as geodesic length $L \geq 0$.

$${}^2\nu \equiv \Phi_b/\hbar, s \equiv \Phi_h/\hbar$$

As $\nu \gg 1$, $L_{\text{ren}} = L - 2 \log(\nu)$ it reduces to $\psi_s(L) \rightarrow K_{2is} (2e^{-L_{\text{ren}}/2})$

Near $L = 0$, actually

$$\psi_s(L) \stackrel{L \rightarrow 0}{\sim} \sqrt{L} + \alpha \sqrt{L} \log L.$$

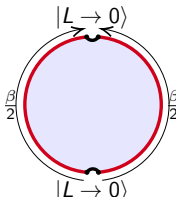
we select $\alpha = 0$ because of embedding into **group theory** .

Orthogonality condition:

$$\int_0^\infty dL \psi_s(L) \psi_{s'}(L) = \frac{\delta(s - s')}{\rho(s)}$$

$$\rho(s) = \frac{s \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)}$$

Partition function \rightarrow **Euclidean transition amplitude** between states of zero geodesic length $|L = 0\rangle$.

$$Z(\beta) = \langle L = 0 | e^{-\beta \hat{H}} | L = 0 \rangle = \frac{\beta}{2} \int_{\frac{\beta}{2}}^{\frac{\beta}{2}} ,$$


The state $|L = 0\rangle$ replaces $|L_{\text{ren}} = -\infty\rangle$!

Partition function

Since for small- L $\psi_s(L) \stackrel{L \rightarrow 0}{\sim} \sqrt{L}$, we extract the leading vanishing behaviour:

$$Z(\beta) = \left(\lim_{L \rightarrow 0} L \right) \int_0^\infty ds \frac{s \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)} e^{-\beta E(s)}.$$

where $E(s) = \nu - \sqrt{\nu^2 - s^2}$.

Issue: what happens when $s > \nu$? Black hole horizon Φ_h becomes larger than cutoff Φ_b . **We will come back to this later.**

Equivalent form

$$Z(\beta) = \left(\lim_{L \rightarrow 0} L \right) \int_0^\infty ds s \sinh(2\pi s) \Gamma\left(\frac{1}{2} \pm is \pm i\nu\right) e^{-\beta E(s)}.$$

Infinite cutoff two-point function with a fixed energy label beyond the boundary particle trajectory and $\Delta = 1/2$.

Comparison with previous work

Our result encodes **two sources of dependence on ν** (Φ_b):

- Deformation of the **Boltzman factor** $\beta E(s) = \beta \left(\nu - \sqrt{\nu^2 - s^2} \right)$

→ at large ν reduces to $\beta_{\text{JT}} s^2$ with $\beta_{\text{JT}} = \beta/\nu$

→ corresponds to the **perturbative branch of $T\bar{T}$ spectrum**³

→ in our case $T\bar{T}$ does not imply $\hat{H}_\lambda = \nu - \sqrt{\nu^2 - (\hat{p}^2 + e^{-\hat{L}_{\text{ren}}})}$

- Deformation of the **spectral density/measure**

$$\rho(s) = \frac{s \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)} \rightarrow s \sinh(2\pi s) + \mathcal{O}(e^{-2\pi \nu})$$

→ **nonperturbative corrections** in the cutoff!

→ consistent with identification parameter μ of **universal cover of $\text{SL}(2, \mathbb{R})$** as $\mu = iB$ ⁴ and $B = 1/\epsilon$ in **JT=particle in a magnetic field**⁵

³Gross, Kruthoff, Rolph, Shaghoulian. Iliesiu, Kruthoff, Turiaci, Verlinde

⁴Iliesiu, Pufu, Verlinde, Wang

⁵Yang, Kitaev

A group theory interpretation

In the infinite cutoff case, using the **Gauss-Euler decomposition** of $SL(2, \mathbb{R})$ group element $g = e^{\gamma F} e^{\phi H} e^{\beta E}$ ⁶

→ fix eigenvalues of **mixed parabolic generators** E, F

→ reduced Casimir equation $\langle \eta | g \mathcal{C} | \mu \rangle$ becomes **Liouville quantum mechanics** $-\frac{d^2}{d\phi^2} - \mu\eta e^{-2\phi}$

If instead we use **KAK decomposition** $g = e^{\theta_1(E-F)} e^{LH} e^{\theta_2(E-F)}$

→ fix eigenvalues of **elliptic generator** $E - F$

→ reduced Casimir operator $\langle m | g \mathcal{C} | n \rangle$ becomes

$$-\frac{d^2}{dL^2} + V_{m,n}(L) \quad V_{m,n}(\phi) = \frac{(m-n)^2 - 1}{4 \sinh^2 L \cosh^2 L} - \frac{mn}{\cosh^2 L}$$

Matches with **quantum Pöschl–Teller Hamiltonian** for $m = n = i\nu = i\Phi_b/\hbar!$

⁶Mertens, Blommaert, Verschelde

- Renormalizing the length $L_{\text{ren}} = L - 2 \log(\Phi_b)$ and using commutation relations

$$e^{\theta_1(E-F)} e^{\frac{1}{2} \log(\phi_b) H} e^{\phi H} e^{\frac{1}{2} \log(\phi_b) H} e^{\theta_2(E-F)}$$

The **KAK decomposition** reduces to **Gauss-Euler** as $\Phi_b \gg 1$.

- **Representation matrix element** (with $\lambda = \frac{1}{2} + is$ principal-series)

$$R_{m,m}(g) \propto e^{im(\theta_1+\theta_2)} (1-u)^\lambda {}_2F_1(\lambda+m, \lambda-m; 1; u), \quad u = \tanh^2 L/2.$$

corresponds to the **bulk wavefunction** written above!

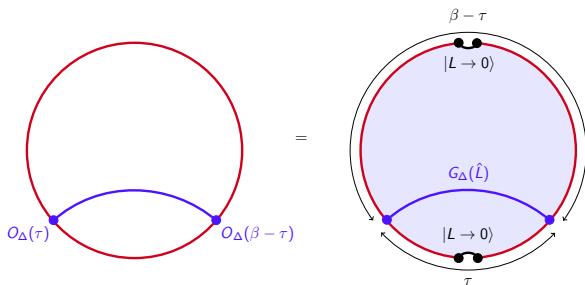
- Analytically continuing representation label, **discrete series** are

$$R_{m,n}^{(\Delta)}(u) = u^{\frac{m-n}{2}} (1-u)^\Delta P_N^{(m-n, 2\Delta-1)}(1-2u)$$

The **lowest-weight coefficient** $\rightarrow R_{0,0}^\Delta \propto \cosh(L/2)^{-2\Delta}$

Analogue of $R_{0,0}^\Delta = e^{-\Delta\phi} \rightarrow$ This should encode **matter insertion!**

Two-point function



For finite cutoff generalization of $e^{-\Delta \hat{L}_{\text{ren}}}$, we **propose insertion of**

$$G_{\Delta}(\hat{L}) = \left[\cosh \left(\frac{\hat{L}}{2} \right) \right]^{-2\Delta} \Phi_b \xrightarrow{\sim \infty} 2^{2\Delta} e^{-\Delta \hat{L}} = \Phi_b^{-2\Delta} e^{-\Delta \hat{L}_{\text{ren}}}.$$

same large-cutoff behavior as the standard exponential proposal. The **matrix element** can be computed exactly!

$$\mathcal{I}_{\cosh}(\Delta; s_1, s_2) := \int_0^{\infty} \psi_{s_1}(L) \psi_{s_2}(L)^* (\cosh(L/2))^{-2\Delta} dL.$$

$$\begin{aligned}
 \langle O_{\Delta}(\tau_1, \tau_2) \rangle_{\beta} &= \text{Diagram} \\
 &= \int_0^{\infty} d\mu(s_1) d\mu(s_2) e^{-(\beta - \tau_2 + \tau_1)E(s_2)} e^{-(\tau_2 - \tau_1)E(s_1)} \times \\
 &\quad \times \Gamma(\Delta \pm is_1 \pm is_2) \mathbb{W}(s_1, \nu; \frac{1}{2} + is_2, \frac{1}{2} - is_2, \Delta - i\nu, \Delta + i\nu)
 \end{aligned}$$

with

$$d\mu(s) \equiv \frac{2s \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi\nu)} ds$$

and $\mathbb{W}(\lambda, x; a, b, c, d)$ is the Wilson function.

→ **encodes the 6j symbol in infinite cutoff JT**

We checked in the **semiclassical** $G_N \rightarrow 0$ **limit**, it localizes to

$$L_E(\tau, \beta) = 2 \sinh^{-1} \left(\frac{\beta}{2\pi} \left| \sin \left(\pi \frac{\tau}{\beta} \right) \right| \right),$$

First finite-cutoff 2pt function proposal to do so!

What happens when $\Phi_h > \Phi_b$?

- $T\bar{T}$ point of view:
 - often associated to resurgence of **nonperturbative branch**⁷
 - not fully clear gravitational interpretation
- **by-hand truncation** of the spectrum at $\Phi_h = \Phi_b$

$$\hat{\Pi} e^{-\beta \hat{H}} \hat{\Pi}, \quad \hat{\Pi} \equiv \int_0^{\Phi_b} s |s\rangle \langle s|.$$

→ state $\hat{\Pi} |L=0\rangle$ is **not a sharply localized** zero-length state anymore

- closer look at the Penrose diagram in global coordinates.

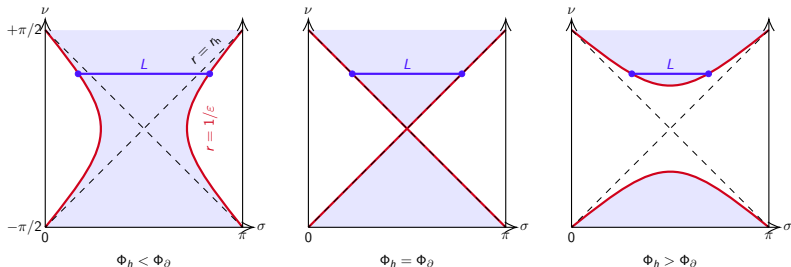
$$ds^2 = \frac{-d\nu^2 + d\sigma^2}{\sin^2(\sigma)}, \quad \Phi = \Phi_h \frac{\cos(\nu)}{\sin(\sigma)},$$

Depending on $\Phi_b, \Phi_h \rightarrow$ **timelike, null or spacelike boundary curves**.
We can still compute the **conserved charge** associated to $\zeta_{L,R}$

$$E(s) = \begin{cases} \Phi_b - \sqrt{\Phi_b^2 - s^2}, & s < \Phi_b, \\ \Phi_b + \sqrt{s^2 - \Phi_b^2}, & s > \Phi_b. \end{cases}$$

⁷Iliesiu, Pufu, Verlinde, Wang. Griguolo, Panerai, Papalini, Seminara

Our proposal: going beyond the horizon



Partition function with **piecewise defined energy function**:

$$Z(\beta) = \int_0^\infty ds \frac{s \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)} e^{-\beta E(s)}.$$

$$E(s) = \begin{cases} \nu^2 - \sqrt{\nu^2 - s^2}, & s < \nu, \\ \nu^2 + \sqrt{s^2 - \nu^2}, & s > \nu. \end{cases}$$

How can we make sense of **space-like evolution** behind the horizon?

→ Possible interpretation in terms of $T\bar{T} + \Lambda$ deformation. ⁸

⁸Ahmad, Almheiri, Lin

Sinh dilaton gravity: summary of main results

- Canonical quantization is performed using **finite cutoff length of ERB in the Weyl rescaled geometry** $ds_{\text{eff}}^2 = e^{2\pi b^2 r} ds^2 \rightarrow$ **effective AdS₂**⁹

- Fine cutoff Hamiltonian has the same $T\bar{T}$ -structure, with

$$H_0 = \sqrt{(1 + e^{4\pi b^2 \Phi_b} e^{-\Lambda})(1 + e^{-\Lambda}) \cosh(\Pi) - e^{2\pi b^2 \Phi_b} e^{-\Lambda}}$$

- Wavefunction is q -hypergeometric function:¹⁰

$$\chi_s(\Lambda) = e^{\Lambda(\frac{1}{2} + i\nu)} F_b \left(\frac{b}{2} + ib\nu + ibs, \frac{b}{2} + ib\nu - ibs, b, -e^{\Lambda - 4\pi b^2 \nu} \right)$$

- Spectral density from orthogonality:

$$\rho(s) = \frac{\sinh(2\pi b^2 s) \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)}$$

- **Reduced $\mathfrak{sl}_q(2, \mathbb{R})$ quantum Casimir** $\langle s_1 | g \mathcal{C} | s_1 \rangle$ with fixed hyperbolic labels $s_1 = \nu$ matches with **finite cutoff Hamiltonian**

⁹Blommaert, Mertens, Papalini.

¹⁰Ponsot, Teschner. **Id**

Outlook

- **First order formulation** of JT gravity at finite cutoff
- Explain the **boundary condition** $E - F$ fixed to be $\propto \Phi_b$ from the bulk
- Physics beyond the horizon in terms of $T\bar{T} + \Lambda$ **deformation**
- Construct finite cutoff amplitudes in terms of **infinite cutoff Feynman rules**
- Explain the **insertion of** $\cosh(L/2)^{-2\Delta}$ from matter coupled to gravity in the bulk
- **OTOC and Lyapunov exponent** at finite cutoff
- **Sine dilaton at finite cutoff**: complexified contour and deformation of DSSYK

Thanks.

Sinh dilaton gravity on shell

$$S = -\frac{1}{2} \int_{\mathcal{M}} d^2x \sqrt{g} \left(\Phi R + \frac{\sinh(2\pi b^2 \Phi)}{\pi b^2} \right) - \oint_{\partial \mathcal{M}} d\tau \sqrt{h} (\Phi K - \text{c.t.})$$

In the gauge $\Phi = r$, classical solution is:

$$ds^2 = - [\cosh 2\pi b^2 r - \cosh 2\pi b^2 \Phi_h] dt^2 + \frac{dr^2}{[\cosh 2\pi b^2 r - \cosh 2\pi b^2 \Phi_h]}$$

Black hole horizon at $r = \Phi_h$ and **holographic screen** at $r = \Phi_b \rightarrow \infty$, where holographic singularity is found \rightarrow not standard holographic setup

Quasi-local energy can be computed again

$$H(\Phi_h) = \sqrt{\cosh(2\pi b^2 \Phi_b)} - \sqrt{\cosh(2\pi b^2 \Phi_b) - \cosh(2\pi b^2 \Phi_h)}.$$

Infinite cutoff lesson: canonical quantization can be performed in an effective Weyl rescaled AdS_2 , corresponding to the saddle of the Liouville field φ ¹¹

Natural: finite cutoff quantization in the effective geometry!

¹¹Blommaert, Mertens, Papalini.

Weyl rescaled geometry

Effective geometry with **same entropy** and **Hawking temperature**:

$$ds_{\text{eff}}^2 = e^{2\pi b^2 r} ds^2 \quad \rho(r) = e^{2\pi b^2 r} - \cosh(2\pi b^2 \Phi_h)$$

Takes the original metric to an **effective AdS₂ black hole**:

$$ds_{\text{eff}}^2 = -(\rho^2 - \rho_h^2(\Phi_h)) dt^2 + \frac{d\rho^2}{\rho^2 - \rho_h^2(\Phi_h)} \quad \rho_h(\Phi_h) = \sinh(2\pi b^2 \Phi_h).$$

Fix the boundary location $r = \Phi_b$ in the original geometry

→ sets a boundary location at $\rho = \rho_b$ **in the effective geometry**:

$$\rho_b = \rho_b(\Phi_b, \Phi_h) = e^{2\pi b^2 \Phi_b} - \cosh(2\pi b^2 \Phi_h).$$

Using the maps $\rho_b(\Phi_b, \Phi_h)$ and $\rho_h(\Phi_h)$, compute **finite cutoff ERB length** in the effective AdS₂:

$$\sinh\left(\frac{L}{2}\right) = \sqrt{\left(\frac{e^{2\pi b^2 \Phi_b} - \cosh(2\pi b^2 \Phi_h)}{\sinh(2\pi b^2 \Phi_h)}\right)^2 - 1} \cosh(t \sinh(2\pi b^2 \Phi_h))$$

Again using $\omega = d\tau \wedge dH = dL \wedge dP$, one finds for $H_0 = \cosh(2\pi b^2 s)$:

$$H_0 = \sqrt{1 + \frac{e^{4\pi b^2 \Phi_b}}{\sinh^2(L/2)} \frac{1}{\tanh(L/2)}} \cosh(2\pi b^2 P \tanh(L/2)) - \frac{e^{2\pi b^2 \Phi_b}}{\sinh^2(L/2)}.$$

Double deformation (b, Φ_b) of Liouville Hamiltonian of JT gravity!

As $b \ll 1 \rightarrow$ **Pöschl–Teller Hamiltonian**

$$\cosh(2\pi b^2 \Phi_h) = 1 + \frac{1}{2} (2\pi b^2)^2 \left(P^2 + \frac{\Phi_b^2}{\cosh^2(L/2)} \right) + \dots$$

As $\Phi_b \gg 1$, $L_{\text{ren}} = L - 4\pi b^2 \Phi_b \rightarrow$ **Liouville gravity Hamiltonian**

$$\lim_{\Phi_b \rightarrow \infty} \cosh(2\pi b^2 \Phi_h) = \sqrt{1 + e^{-L_{\text{ren}}}} \cosh(2\pi b^2 P).$$

Perform a canonical transformation:

$$\Lambda := 2 \ln \left(\sinh \frac{L}{2} \right), \quad \Pi := P \tanh \frac{L}{2},$$

More pleasant form:

$$H_0 = \sqrt{(1 + e^{4\pi b^2 \Phi_b} e^{-\Lambda}) (1 + e^{-\Lambda})} \cosh(\Pi) - e^{2\pi b^2 \Phi_b} e^{-\Lambda}$$

Schrödinger equation and wavefunction

$$\begin{aligned} & \frac{1}{2} \left(1 + e^{-\Lambda + i\pi b^2} \right) \chi_s(\Lambda - 2\pi i b^2) + \frac{1}{2} \left(1 + e^{4\pi b^2 \nu} e^{-\Lambda - i\pi b^2} \right) \chi_s(\Lambda + 2\pi i b^2) \\ & = \left(\cosh(2\pi b^2 s) + e^{2\pi b^2 \nu} e^{-\Lambda} \right) \chi_s(\Lambda). \end{aligned}$$

As $b \rightarrow 0$, it yields the quantum Pöschl–Teller potential!

In Fourier space:

$$\frac{\tilde{\chi}_s(k+i)}{\tilde{\chi}_s(k)} = e^{-2\pi b^2 \nu} \frac{\sinh(\pi b^2(s+k)) \sinh(\pi b^2(s-k))}{\sinh(\pi b^2(k+\nu+i/2))^2}.$$

A solution can be found

$$\chi_s(\Lambda) = \frac{S_b(b)}{S_b(\frac{b}{2} + ib\nu \pm ibs)} \int dk e^{-i\Lambda k} \frac{S_b(-ibk \pm ibs) S_b(\frac{b}{2} + ib\nu + ibk)}{S_b(\frac{b}{2} - ib\nu - ibk)} e^{2\pi i b^2 \nu k}$$

→ Mellin-Barnes representation of **q -hypergeometric function**:¹²

$$\chi_s(\Lambda) = e^{\Lambda(\frac{1}{2} + i\nu)} F_b \left(\frac{b}{2} + ib\nu + ibs, \frac{b}{2} + ib\nu - ibs, b, -e^{\Lambda - 4\pi b^2 \nu} \right)$$

¹²Ponsot, Tschner, Ip

All the **correct limits** are verified:

As $\nu \rightarrow \infty$, with $\Lambda_{\text{ren}} = \Lambda - 4\pi b^2 \nu$ it reduces to q -deformed BesselK, i.e.

Liouville gravity wavefunction¹³

$$\rightarrow e^{\pi\nu} \int_{-\infty}^{+\infty} dk e^{-i\Lambda_{\text{ren}}k} S_b(-ibk \pm ibs) e^{i\pi b^2(k^2 - s^2)}$$

As $b \rightarrow 0$, using $S_b(bx) \rightarrow \Gamma(x)$ it reduces to **standard JT hypergeometric**

$$\rightarrow \tanh(L/2) (\cosh(L/2))^{1+2i\nu} {}_2F_1\left(\frac{1}{2} + i\nu + is, \frac{1}{2} + i\nu - is, 1; -\sinh^2(L/2)\right)$$

WKB scattering region: encoded by poles at $k = \pm s$

$$\chi_s(\Lambda) = A_s e^{-i\Lambda s} (1 + \mathcal{O}(e^{-\Lambda})) + A_{-s} e^{i\Lambda s} (1 + \mathcal{O}(e^{-\Lambda})) .$$

Asymptotic structure of $\chi_s(\Lambda)$ as $\Lambda \rightarrow \infty$, with

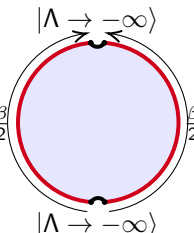
$$A_s = \frac{S_b(-2ibs)}{S_b\left(\frac{b}{2} - ib\nu - ibs\right) S_b\left(\frac{b}{2} + ib\nu - ibs\right)} e^{2\pi i b^2 \nu s} .$$

¹³Mertens, Turiaci, Fan

Spectral density and partition function

We can infer the orthogonality condition $\int d\Lambda \chi_s^*(\Lambda) \chi_{s'}(\Lambda) = \frac{\delta(s-s')}{\rho(s)}$

$$\rho(s) = |A_s|^2 = \frac{\sinh(2\pi b^2 s) \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)}$$



$$Z(\beta) = \frac{\beta}{2} \int_{\beta/2}^{\beta/2} = \left(\lim_{\Lambda \rightarrow -\infty} e^\Lambda \right) \int_0^{+\infty} ds |\psi_s(\Lambda)|^2 e^{-\beta E(s)},$$

$$Z(\beta) = \int_0^\infty ds \frac{\sinh(2\pi b^2 s) \sinh(2\pi s)}{\cosh(2\pi s) + \cosh(2\pi \nu)} e^{-\beta E(s)}.$$

$$E(s) = \begin{cases} \cosh(2\pi b^2 \nu) - \sqrt{\cosh(2\pi b^2 \nu) - \cosh(2\pi b^2 s)}, & s < \nu, \\ \cosh(2\pi b^2 \nu) + \sqrt{\cosh(2\pi b^2 s) - \cosh(2\pi b^2 \nu)}, & s > \nu. \end{cases}$$

A quantum group perspective

Standard Liouville gravity \rightarrow **mixed-parabolic Casimir equation** \mathcal{C}
(where $T_a^\phi f(\phi) = f(\phi + a)$)

$$\frac{1}{2} \left(T_{i\pi b^2}^\phi + T_{-i\pi b^2}^\phi + \mu\eta e^{-2\phi} \right) \psi_{\mu\eta}(g) = \cosh(2\pi bs) \psi_{\mu\eta}(g)$$

and its $b \rightarrow 1/b$ dual $\tilde{\mathcal{C}}$. Unique solution is **Whittaker function**:

$$\psi_{\mu\eta}(g) = \int_{\mathcal{C}} d\xi \mu^{\frac{1}{b}(i\xi - is)} \eta^{\frac{1}{b}(i\xi + is)} S_b(-i\xi \pm is) e^{-2i\phi\xi/b}$$

From **mixed-parabolic to hyperbolic** matrix element

$$\langle \eta | g | \mu \rangle = \int_{-\infty}^{\infty} ds_1 ds_2 \langle \eta | s_1 \rangle \langle s_1 | g | s_2 \rangle \langle s_2 | \mu \rangle$$

Exploiting the Gauss-Euler decomposition $g = g_b(\gamma F) e^{2\phi H} g_b^*(\beta E)$: ¹⁴

$$R_{s_1 s_2}(g) = \frac{1}{\mathcal{N}} \int_0^{+\infty} d\mu \mu^{is_2 + is/b - \frac{1}{2}} g_b^*(\beta\mu) \int_{\mathcal{H}} d\eta \eta^{is_1 - is/b - \frac{1}{2}} g_b(-\gamma\eta) \psi_{\mu\eta}(\phi)$$

The hyperbolic matrix element $R_{s_1 s_2}(g) \equiv \langle s_1 | g | s_2 \rangle$ and the Whittaker function related by **double integral transform**.

¹⁴Mertens

Acting with **double transform** $\mathcal{L}_\mu \mathcal{L}_\eta$ on the Casimir equation $\langle \eta | g \mathcal{C} | \mu \rangle$

→ **Reduced Casimir** $\langle s_1 | g \mathcal{C} | s_1 \rangle$ on hyperbolic matrix elements

$$\begin{aligned} & \frac{1}{2} \left(T_{i\pi b^2}^\phi + T_{-i\pi b^2}^\phi + (\gamma e^{2\phi} \beta)^{-1} (e^{i\pi b^2} e^{-2\pi b s_1} T_{-i\pi b^2}^\phi + e^{-i\pi b^2} e^{2\pi b s_1} T_{i\pi b^2}^\phi - 2) \right) R_{s_1 s_1}(g) \\ &= \cosh(2\pi b s) R_{s_1 s_1}(g) \end{aligned}$$

If we identify hyperbolic element as

$$\gamma e^{2\phi} \beta \equiv e^{\Lambda - 2\pi b^2 \nu}$$

→ **Finite cutoff Hamiltonian derived from the bulk!**

Finally, performing the double integral transform

$$R_{s_1 s_1}(g) = \frac{\gamma^{-is_1 - 1/2} \beta^{-is_1 - 1/2}}{S_b\left(\frac{b}{2} + ibs_1 \pm is\right)} \int d\xi \left(\gamma e^{2\phi} \beta\right)^{-i\xi/b} S_b(-i\xi \pm is) \frac{S_b^2\left(\frac{b}{2} + ibs_1 + i\xi\right)}{S_b^2\left(\frac{b}{2} - ibs_1 - i\xi\right)}$$

Finite cutoff wavefunction $\chi_s(\Lambda)$ → **hyperbolic representation matrix element** $\langle s_1 | g | s_1 \rangle$ **with hyperbolic label** s_1 → **cutoff!**

Towards sine dilaton gravity

Quasi local energy $H = \sqrt{-2 \cos(\Phi_b)} - \sqrt{-2 \cos(\Phi_b) + 2 \cos(\theta)}$

$$H_0 = 2 \cos(\theta) = e^{-i\hat{\Pi}} + 2e^{-i\Phi_b} e^{-\hat{\Lambda}} + \left(1 - e^{-2i\Phi_b} e^{-\hat{\Lambda}}\right) e^{i\hat{\Pi}} \left(1 - e^{-\hat{\Lambda}}\right)$$

- H and H_0 are both real along an imaginary contour $\Phi_b \in [\pi, \pi + i\infty]$.
- As $\nu = i\Phi_b/\hbar \rightarrow \infty$, $\Lambda_{\text{ren}} = \Lambda + 2\nu$ it reduces to **DSSYK transfer matrix!**

Classical length Λ (after removing a pure complex phase $+i\pi/2$)

$$\Lambda = 2 \log \left(\sqrt{\frac{(e^{-ir_b} - \cos(\theta))^2}{\sin^2(\theta)} + 1} \cosh(\sin(\theta)t) \right)$$

Minimum value allowed in phase space at $(t = 0, \theta = \arccos(e^{i\Phi_b}))$:

$$\Lambda^{\min} = -2\nu \quad \Lambda_{\text{ren}}^{\min} = 0$$

- explains the puzzle $\Lambda_{\text{ren}} = 2|\log q|n \geq 0$ **length positivity constraint!**
- Al-Salam-Chihara polynomials $Q_n(\cos(\theta)|e^\nu, e^\nu; q^2)$ is solution.

More on the choice of ordering

Using the mapping

$$\hat{\Lambda} \equiv 2 \log \left(\sinh \left(\hat{L}/2 \right) \right),$$
$$\hat{\Pi} \equiv \frac{1}{2} \left(\hat{P} \tanh \left(\hat{L}/2 \right) + \tanh \left(\hat{L}/2 \right) \hat{P} \right).$$

The $b \rightarrow 0$ limit of sinh dilaton is

$$1 + \frac{B^2}{2} \hat{\Phi}_h^2 + O(B^4)$$
$$= 1 + \frac{B^2}{2} \left(\frac{1}{4\kappa^2} \hat{\Pi}^2 (1 + e^{-\hat{\Lambda}}) + \frac{1}{4\kappa^2} (1 + e^{-\hat{\Lambda}}) \hat{\Pi}^2 + \frac{1}{2\kappa^2} \hat{\Pi} (1 + e^{-\hat{\Lambda}}) \hat{\Pi} + \frac{\Phi_b^2}{1 + e^{\hat{\Lambda}}} \right)$$

The corresponding JT gravity prepotential operator takes the form

$$\hat{\Phi}_h^2(\hat{L}, \hat{P})$$
$$= \frac{3}{4} \hat{P}^2 + \frac{1}{8} \left(\frac{1}{\tanh(\hat{L}/2)} \hat{P}^2 \tanh(\hat{L}/2) + \tanh(\hat{L}/2) \hat{P}^2 \frac{1}{\tanh(\hat{L}/2)} \right) + \frac{\Phi_b^2}{\cosh^2(\hat{L}/2)},$$

which yields the JT Schroedinger equation and the reduced $s/(2, \mathbb{R})$ Casimir.

$$Z(\beta) = \left(\lim_{L \rightarrow 0} L \right) \int_0^\infty ds \frac{s \sinh(2\pi s)}{\pi^3} \Gamma \left(\frac{3}{4} \pm i\delta \pm is \right) e^{-\beta E(s)}.$$

where

$$s \equiv \frac{\Phi_h}{\hbar}, \quad \frac{\Phi_b^2}{\hbar^2} \equiv \delta^2 + \frac{1}{16}.$$